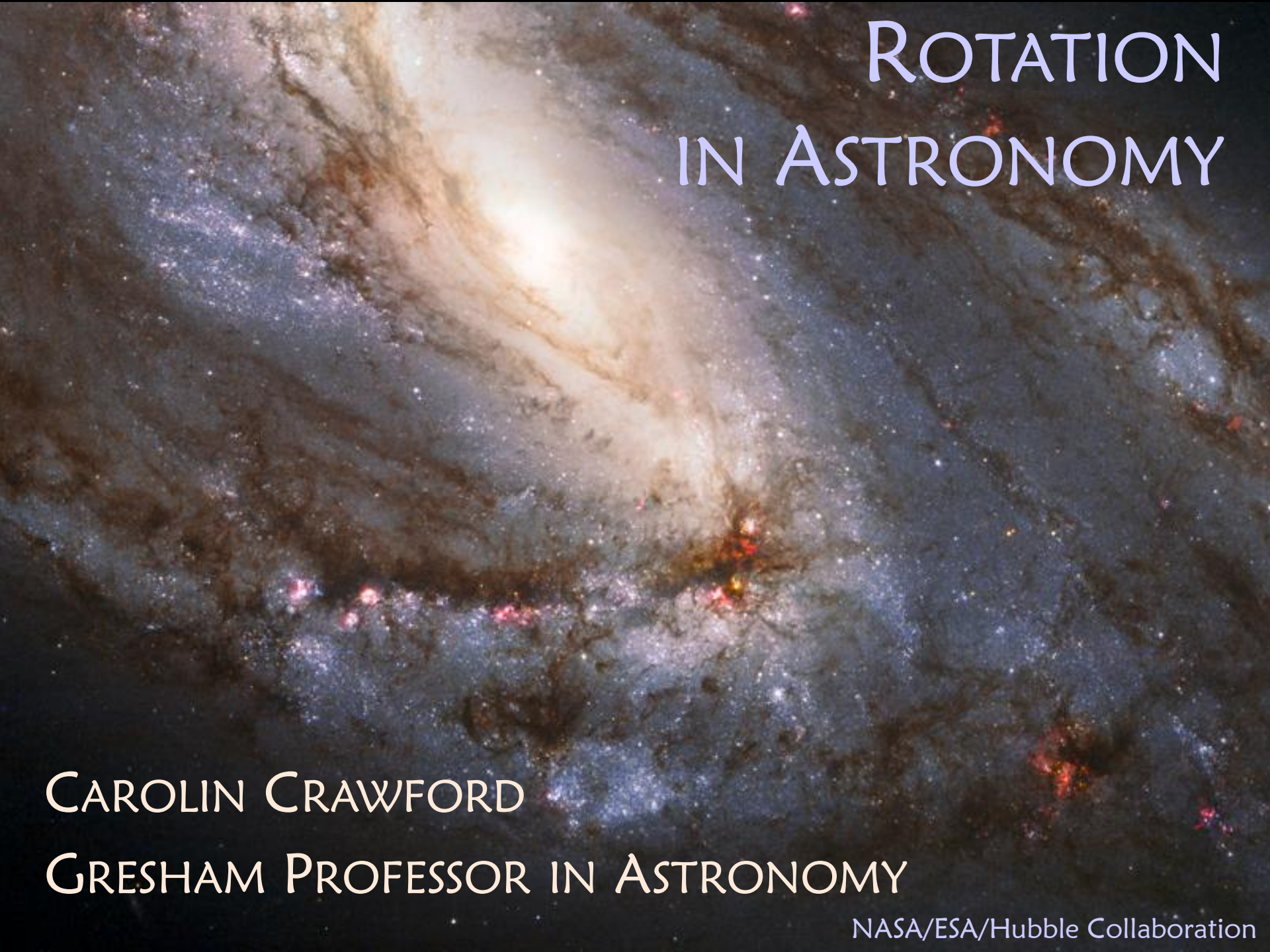




# ROTATION IN ASTRONOMY

CAROLIN CRAWFORD

GRESHAM PROFESSOR IN ASTRONOMY

NASA/ESA/Hubble Collaboration



NASA Suomi NPP mission



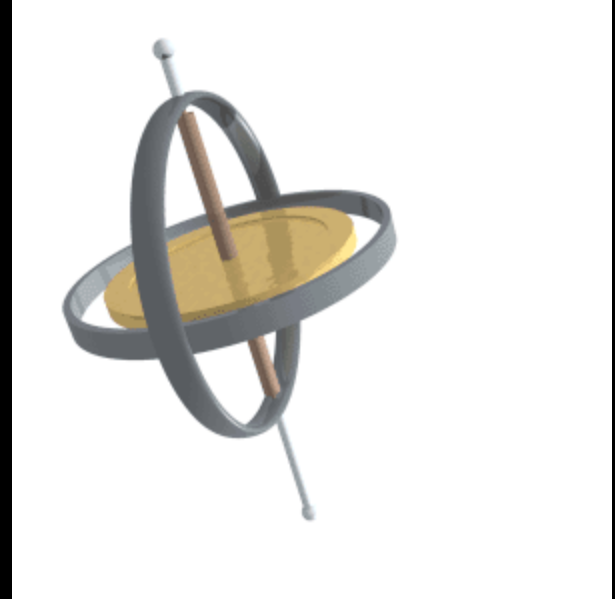
NASA NEAR mission

# ANGULAR MOMENTUM



NASA/Hubble Heritage Team

# ANGULAR MOMENTUM

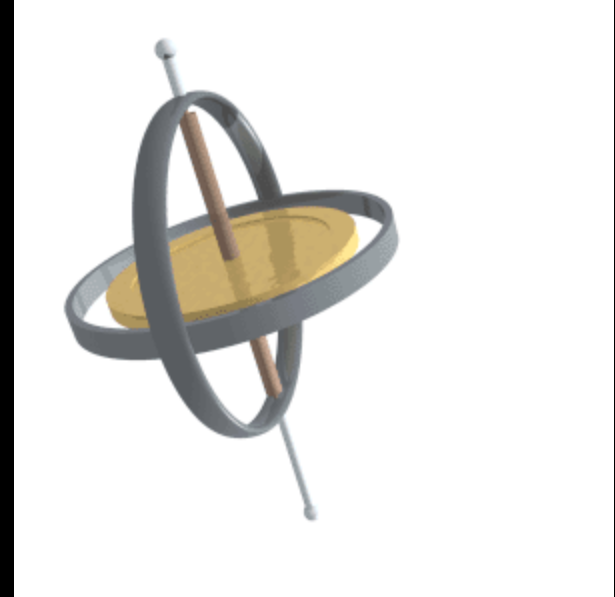


# ANGULAR MOMENTUM

for a rotating body, the quantity

mass x rotational rate x (radius)<sup>2</sup>

is conserved about an axis



# ANGULAR MOMENTUM

for a rotating body, the quantity

mass x rotational rate x (radius)<sup>2</sup>

is *conserved* about an axis

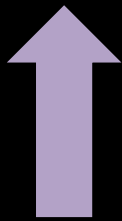
# ANGULAR MOMENTUM

for a rotating body, the quantity

mass x rotational rate x (radius)<sup>2</sup>

is conserved about an axis

if radius

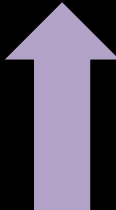


# ANGULAR MOMENTUM

for a rotating body, the quantity

mass  $\times$  rotational rate  $\times$  (radius)<sup>2</sup>

is conserved about an axis

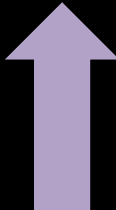
if radius  then rotational rate must 

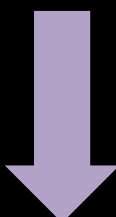
# ANGULAR MOMENTUM

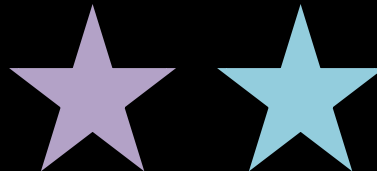
for a rotating body, the quantity

$$\text{mass} \times \text{rotational rate} \times (\text{radius})^2$$

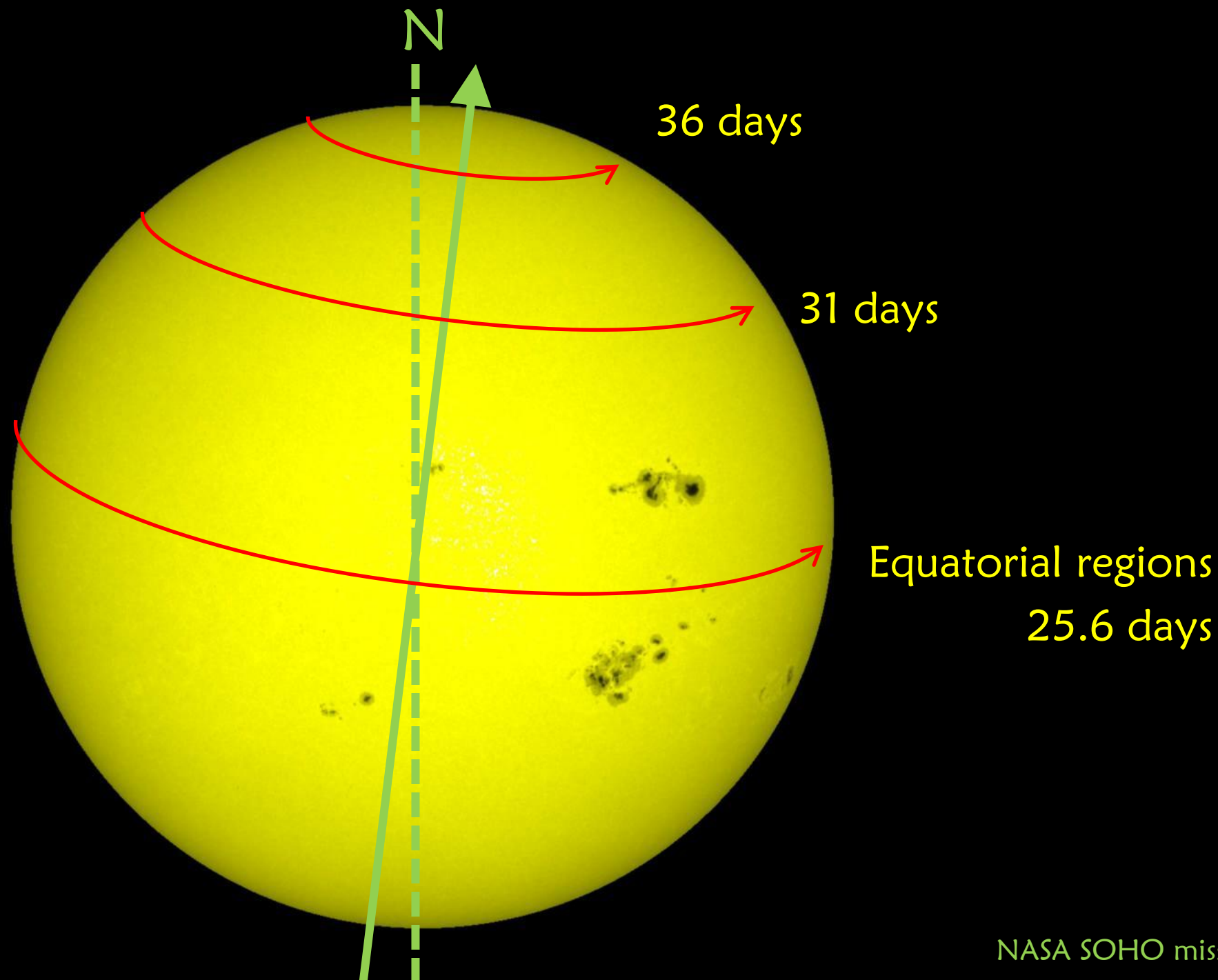
is conserved about an axis

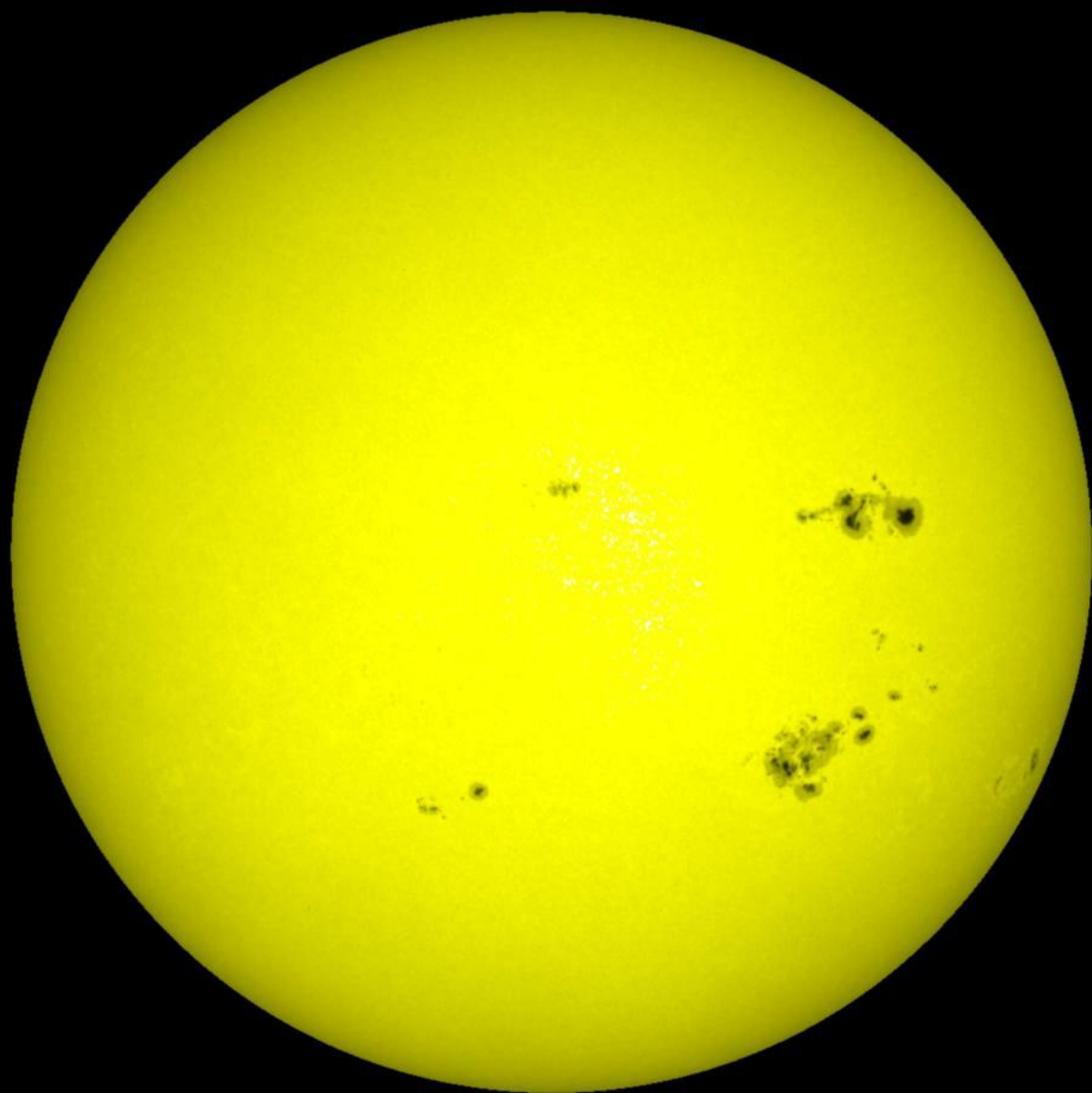
if radius  then rotational rate must 

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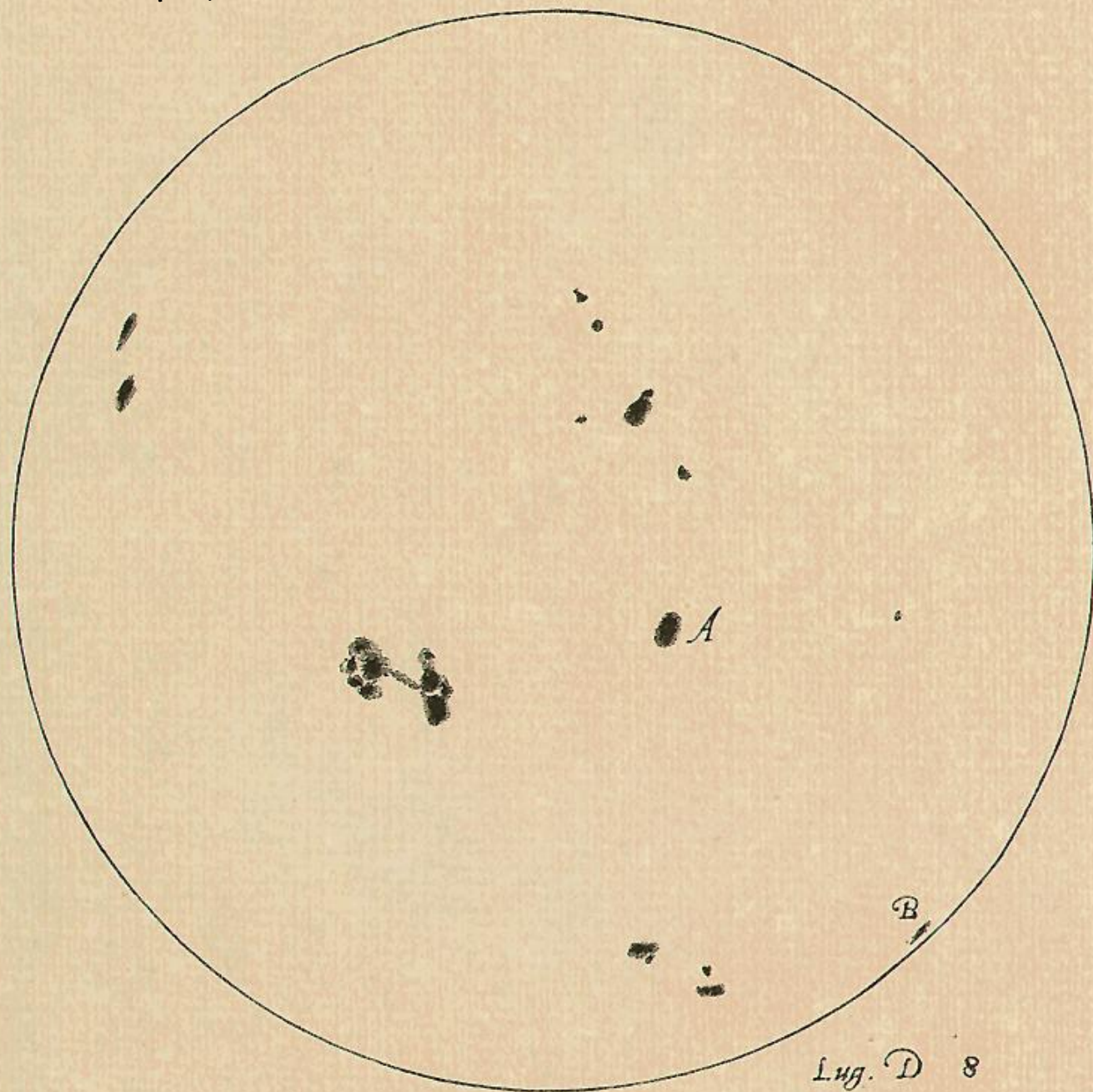








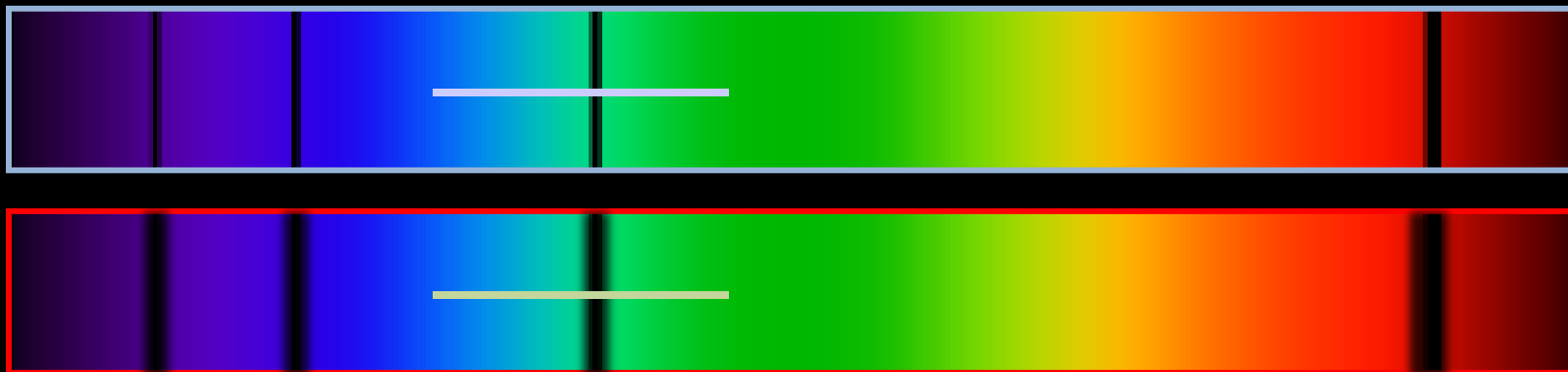
NASA SOHO mission





blueshift

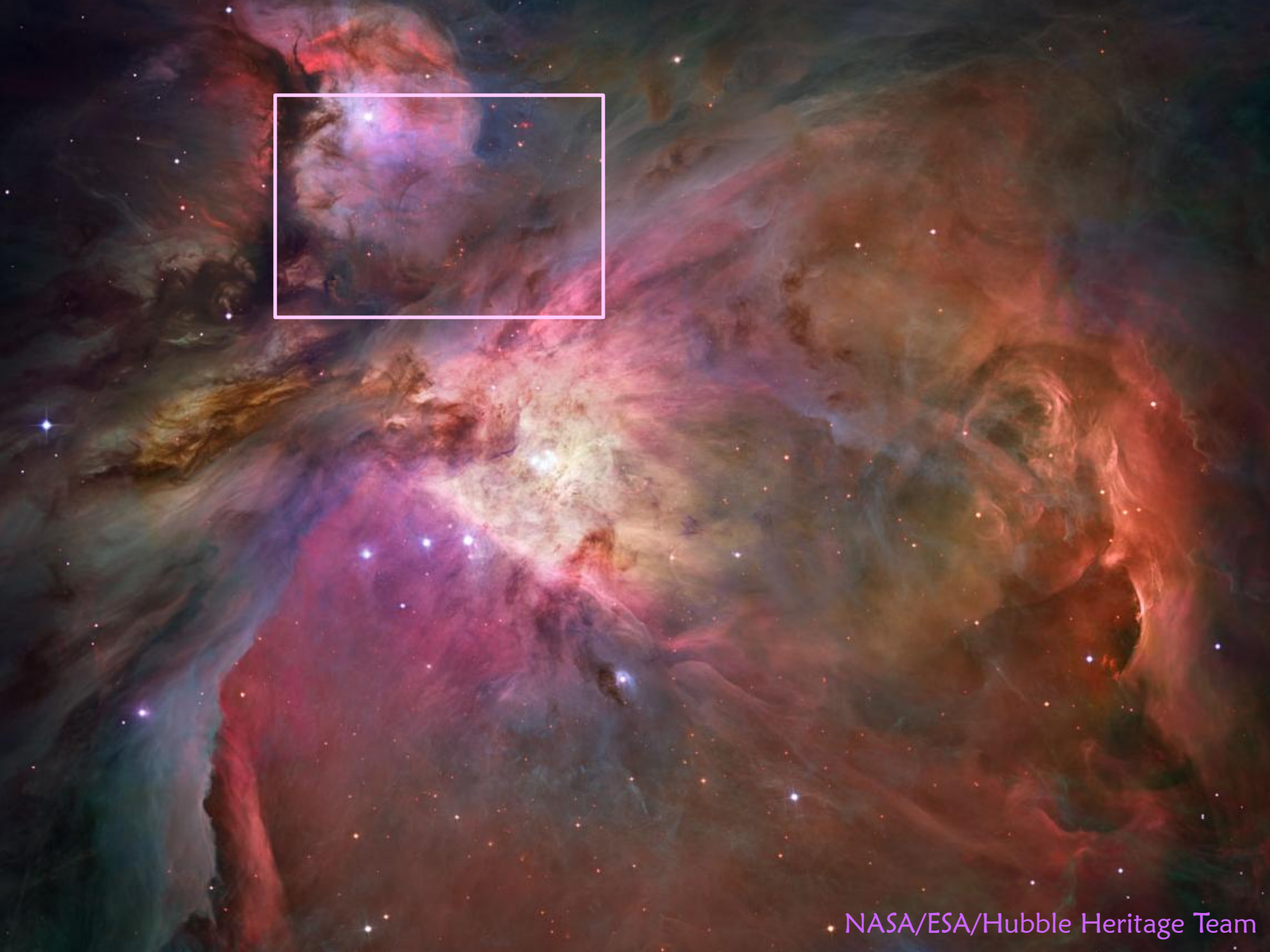
redshift

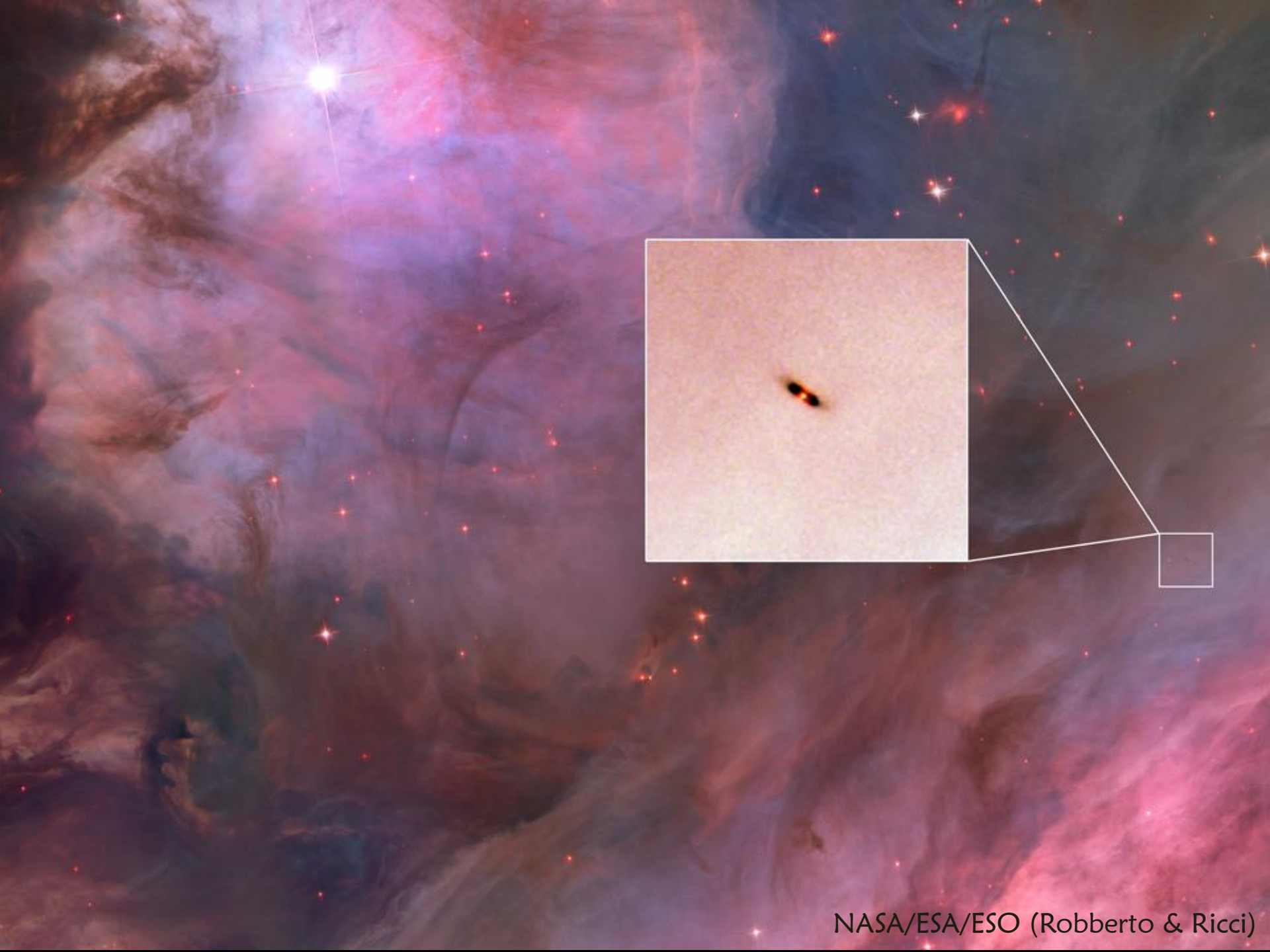


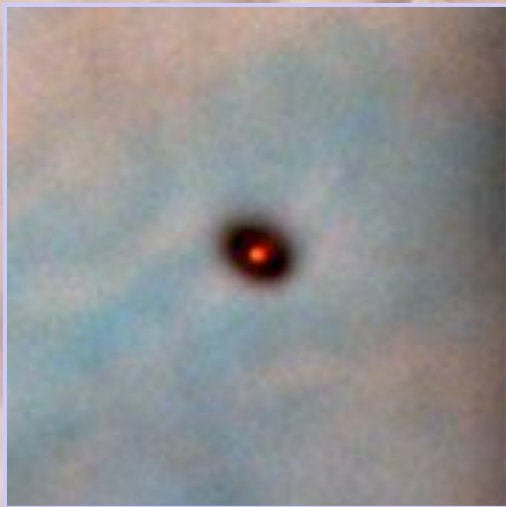


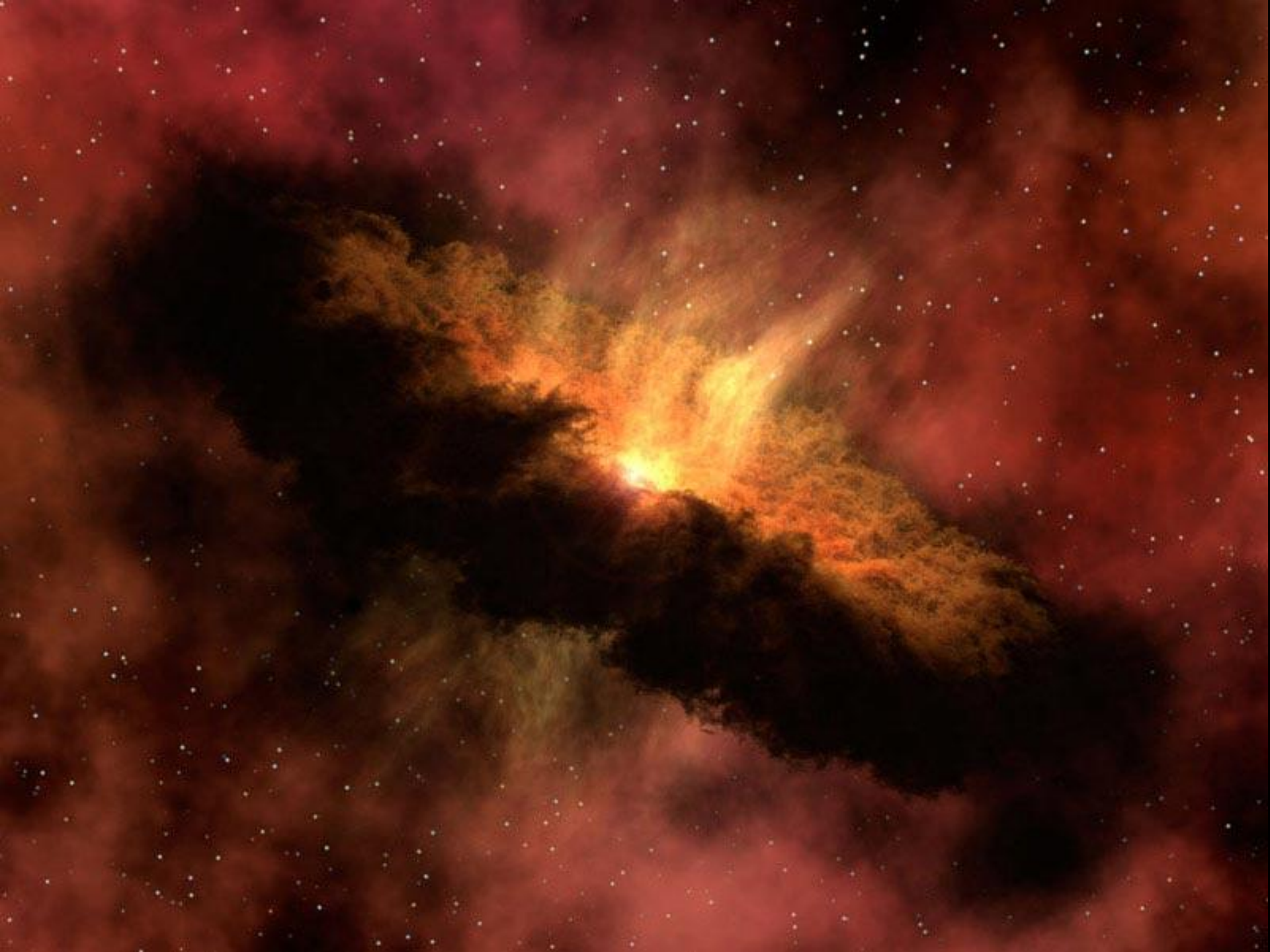


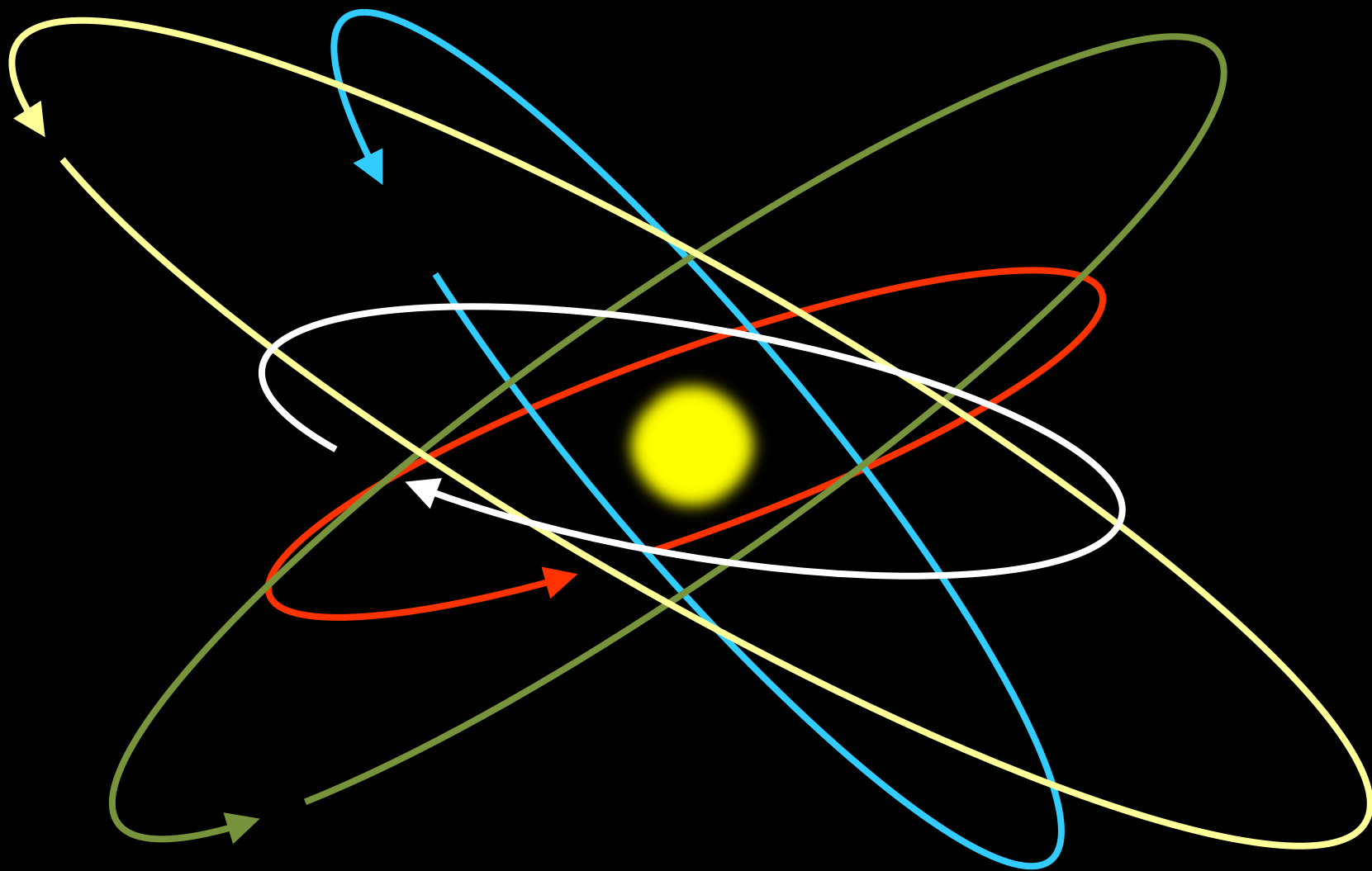
NASA/ESA/Hubble Heritage Team



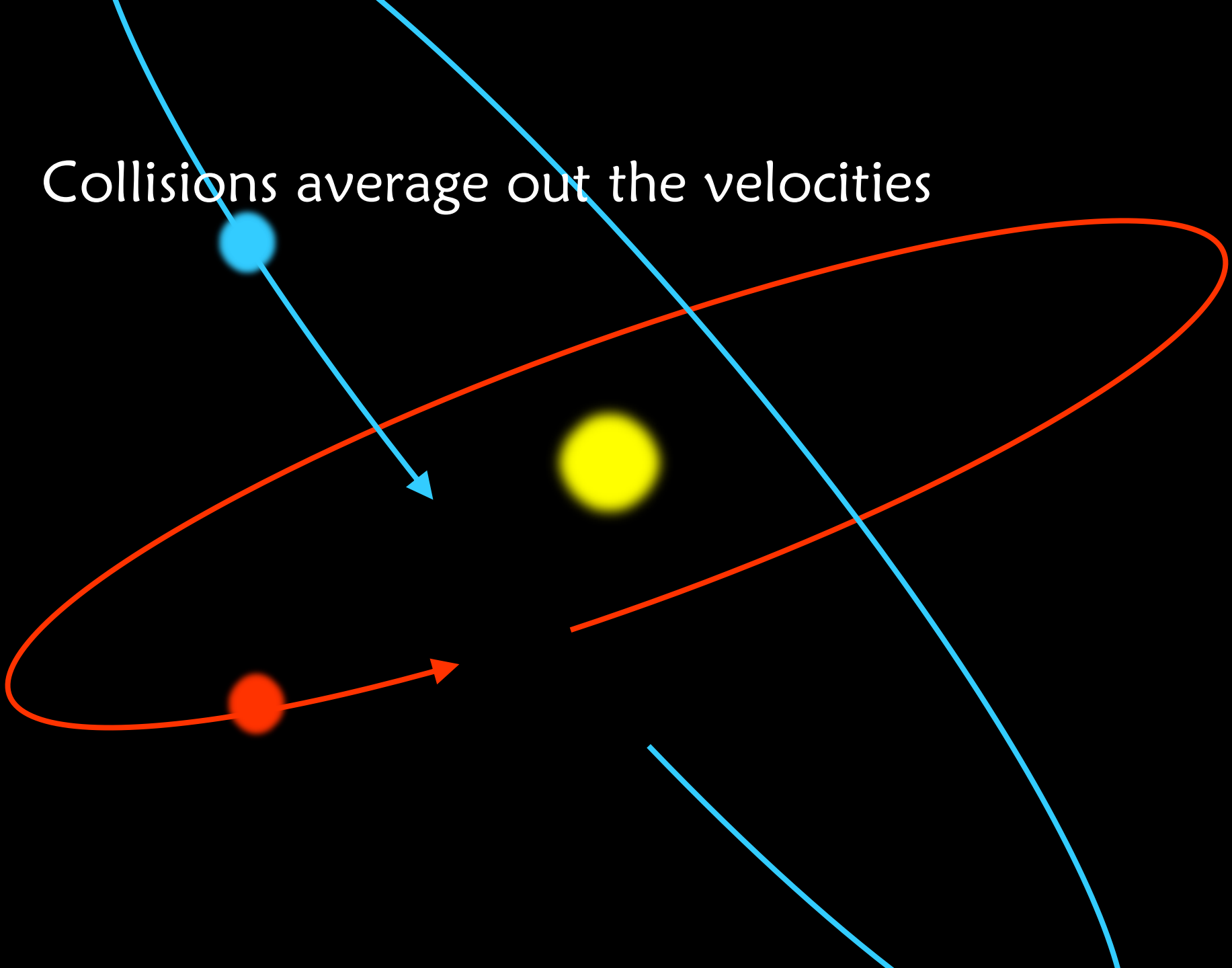




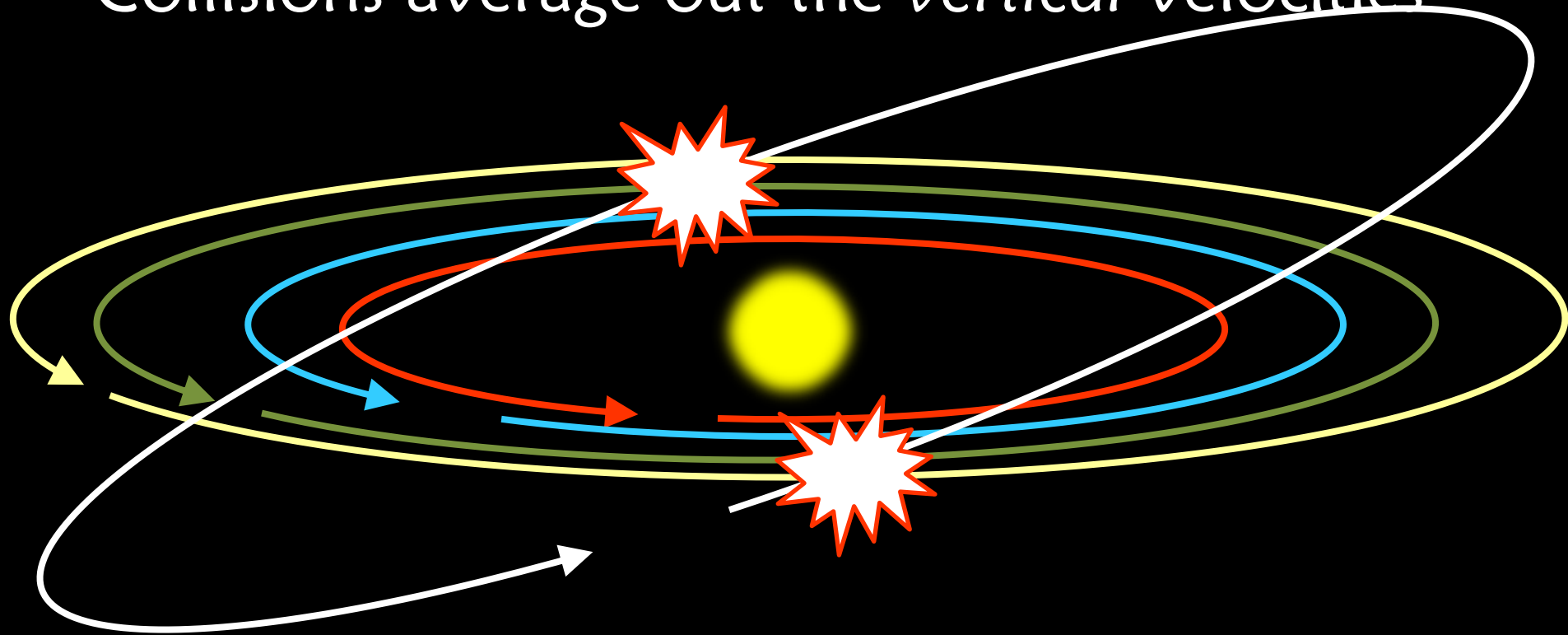


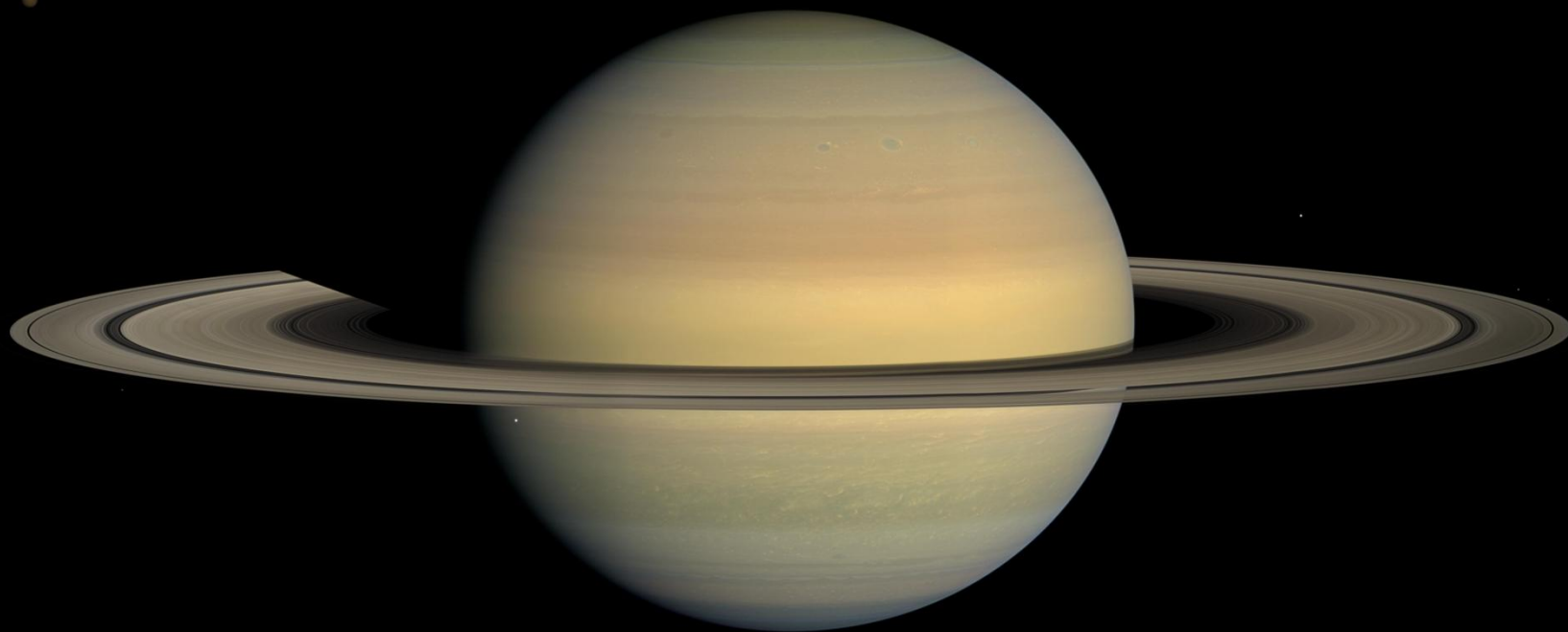


Collisions average out the velocities



Collisions average out the *vertical* velocities

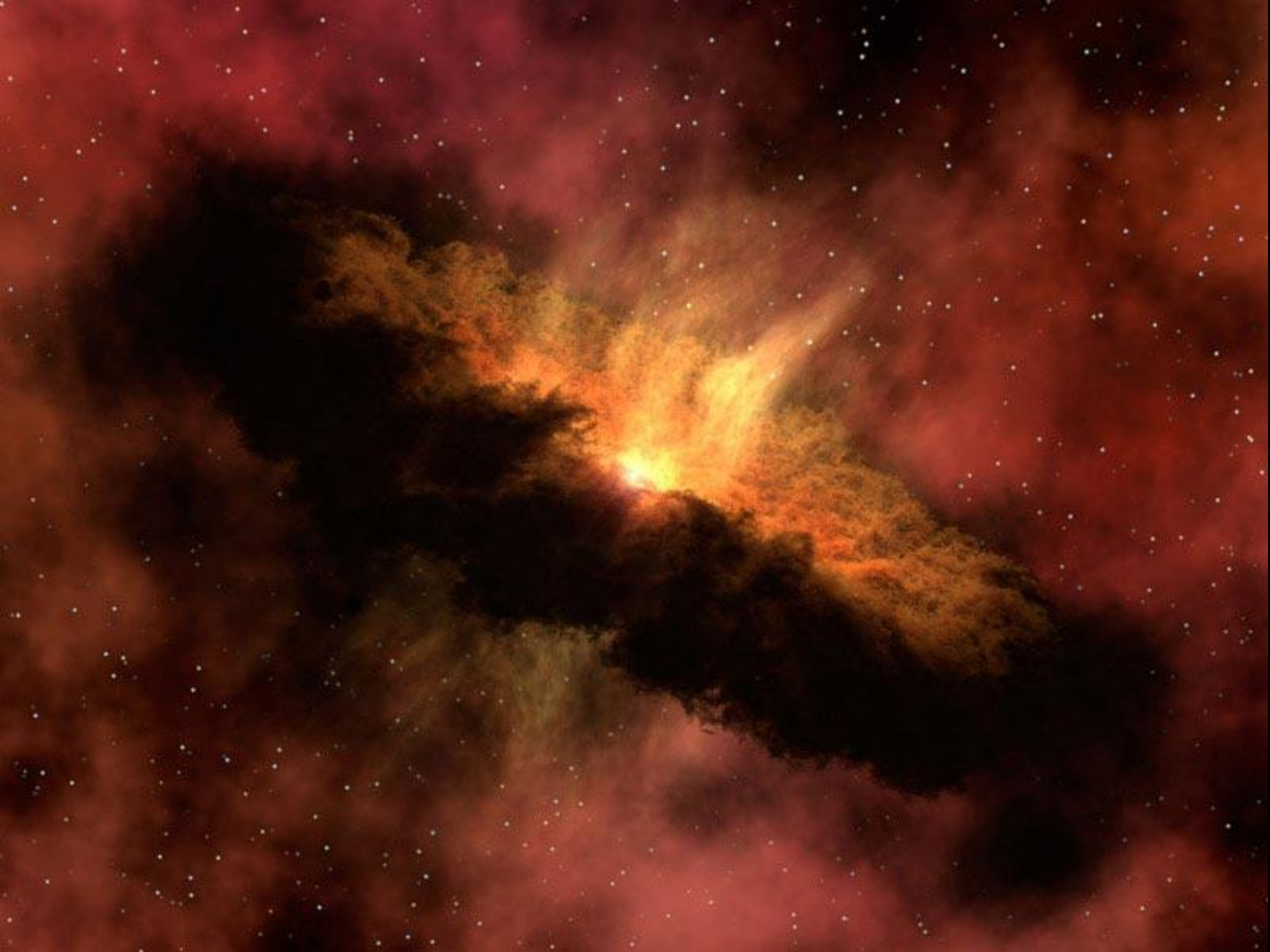


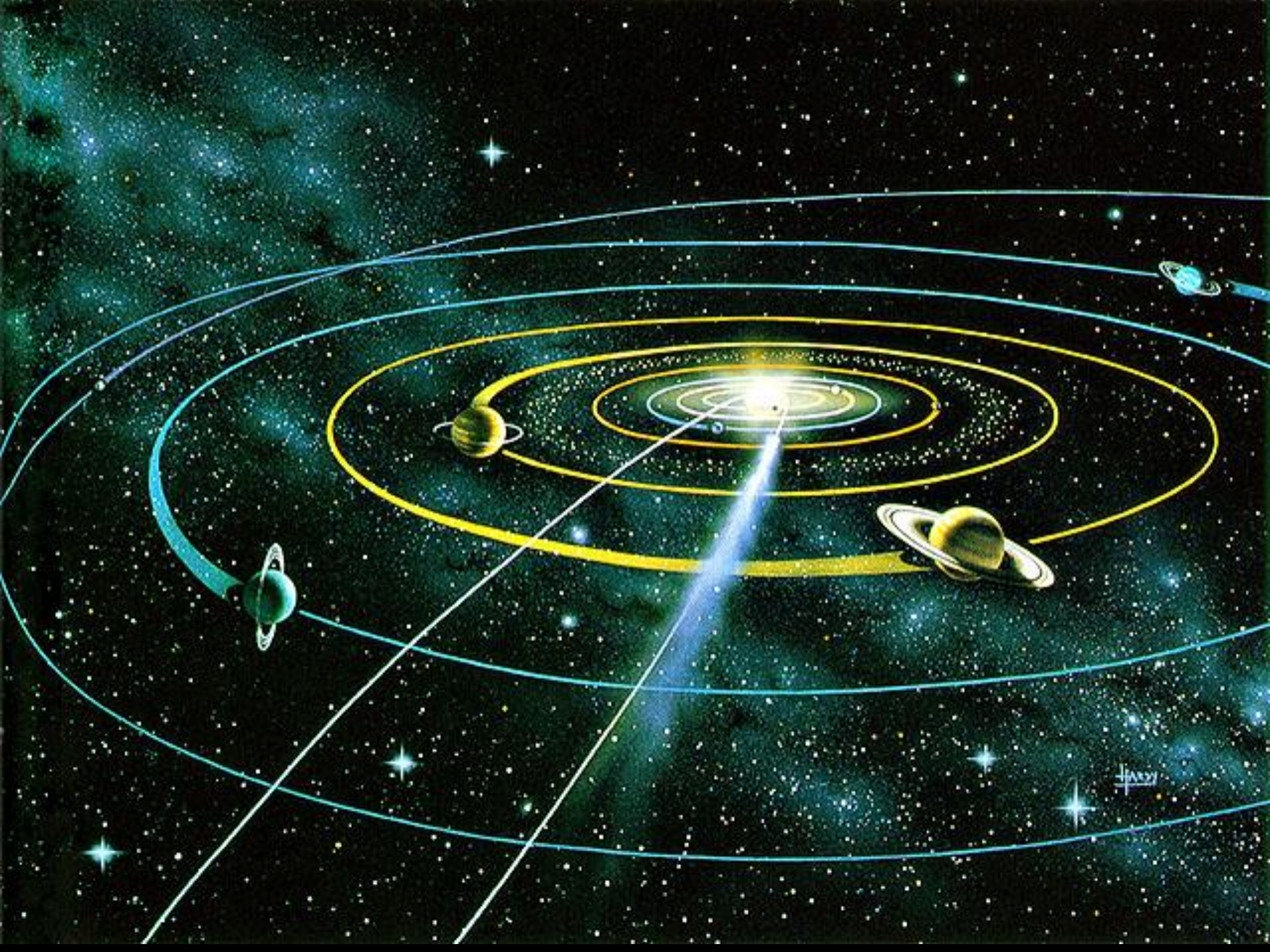


NASA *Cassini* mission

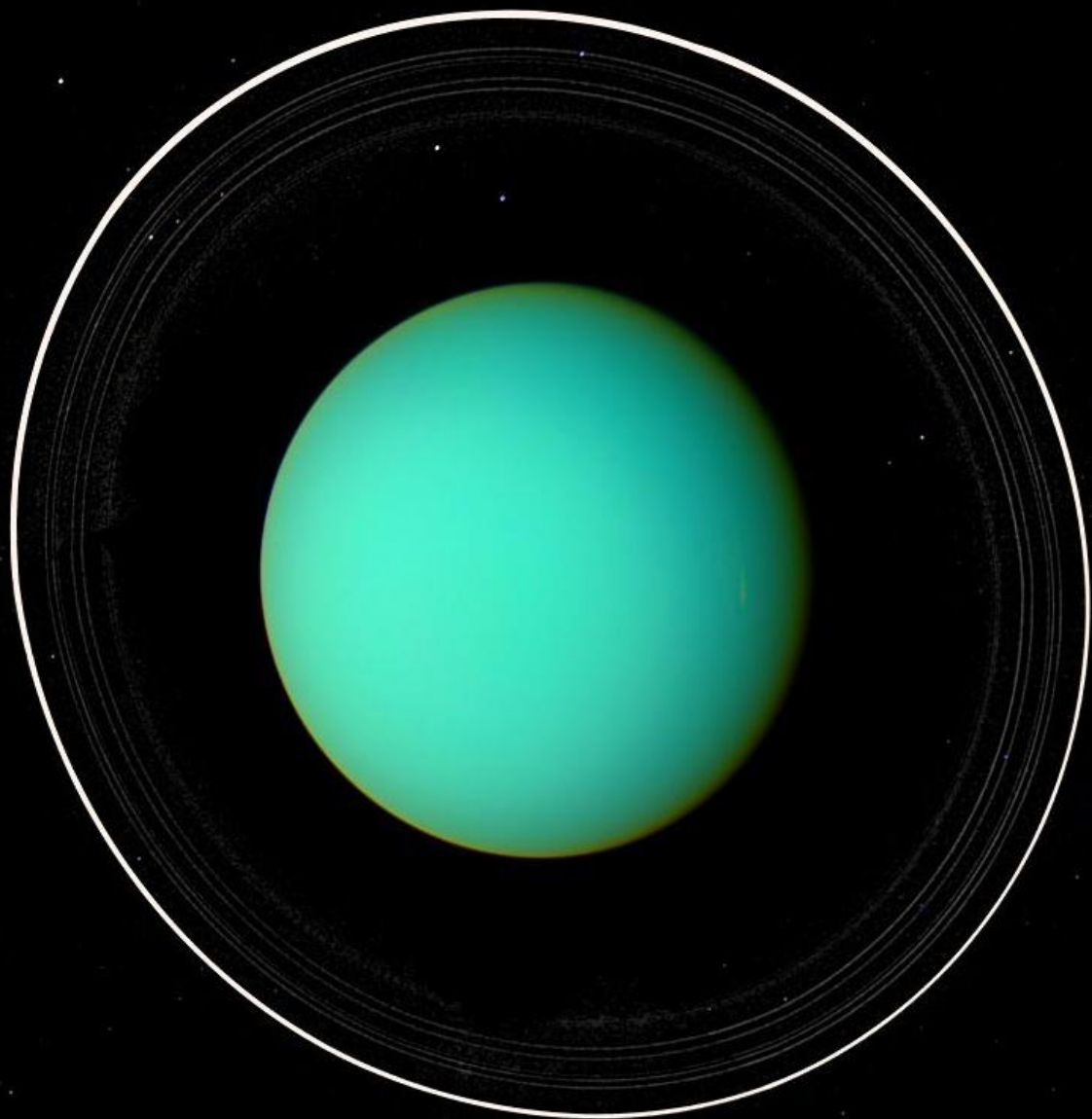


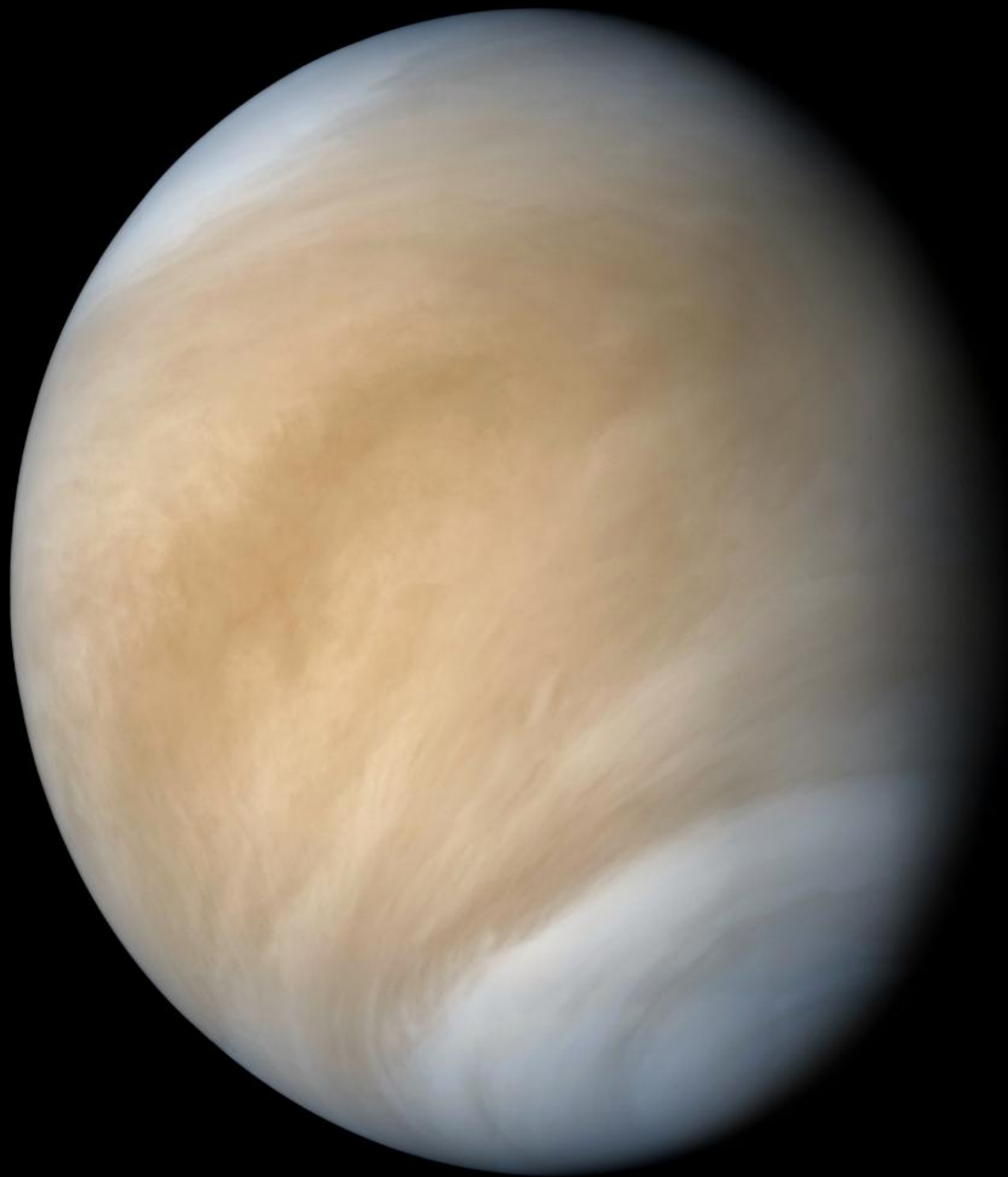










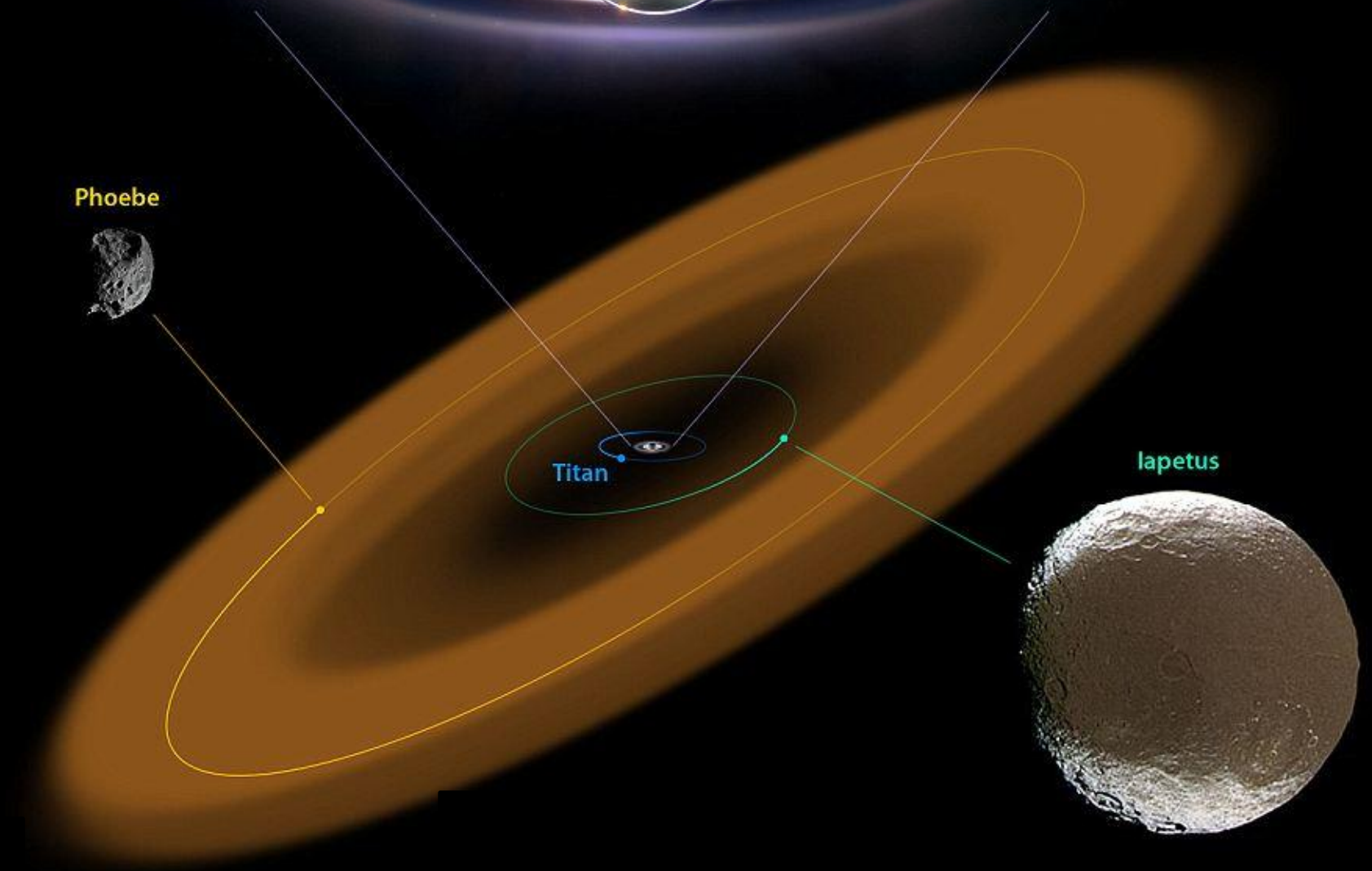


Saturn

Phoebe

Titan

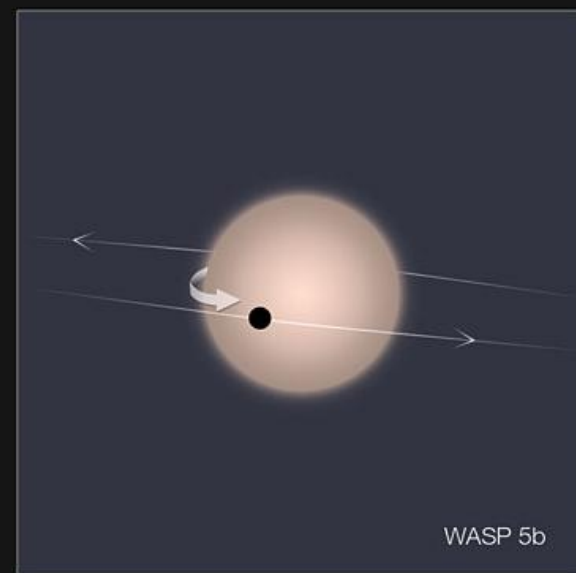
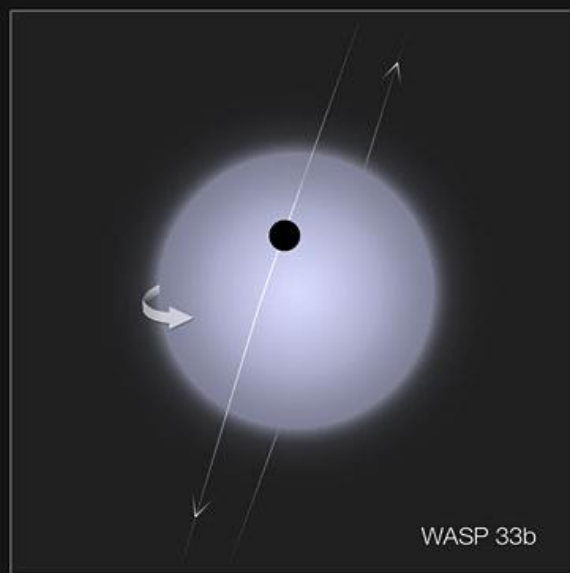
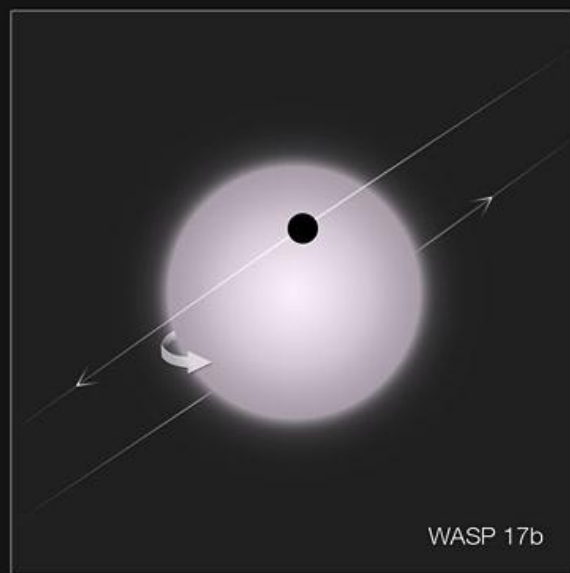
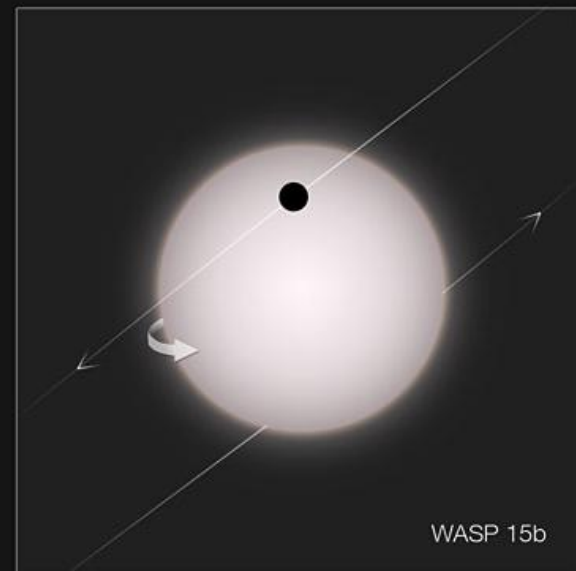
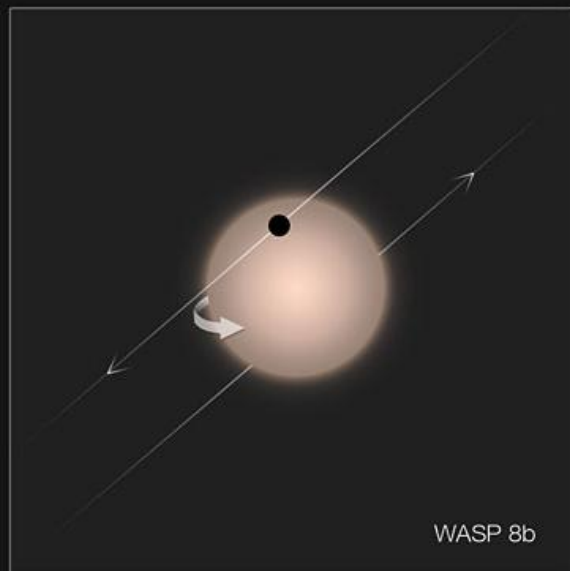
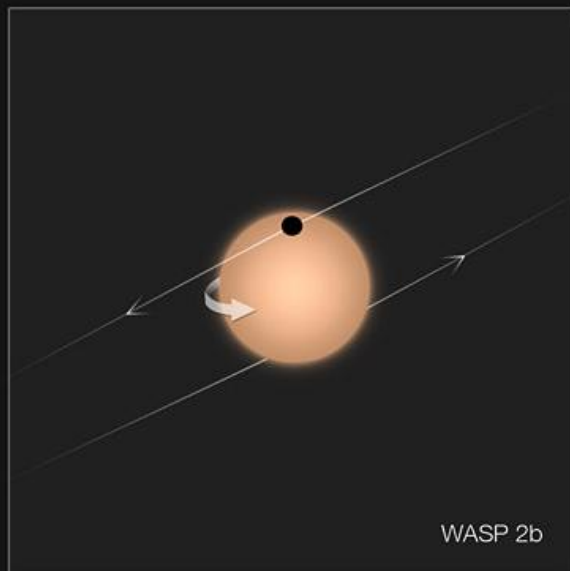
Iapetus





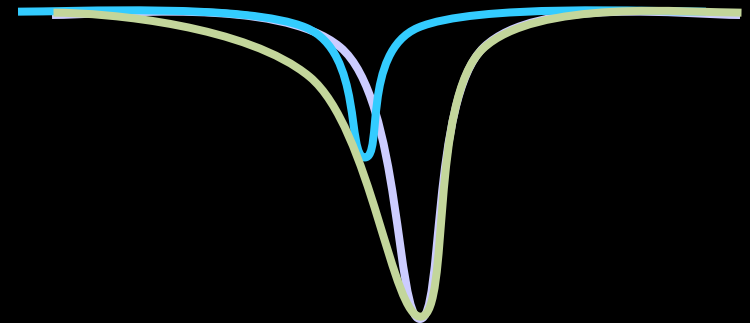
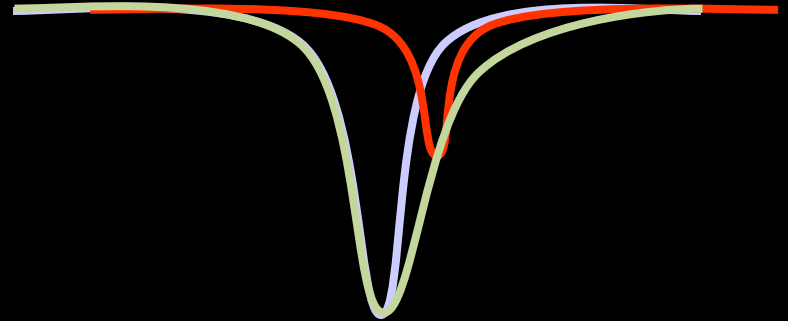
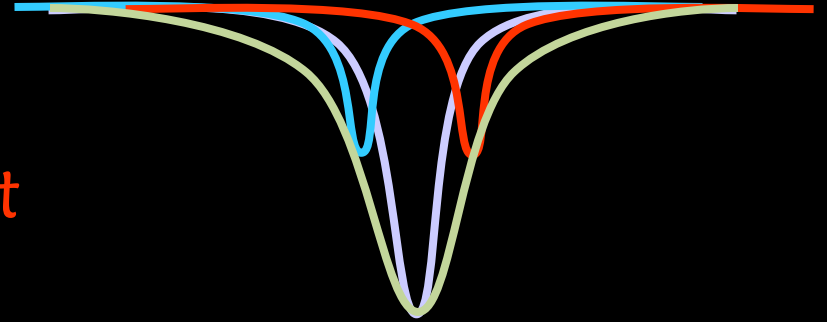
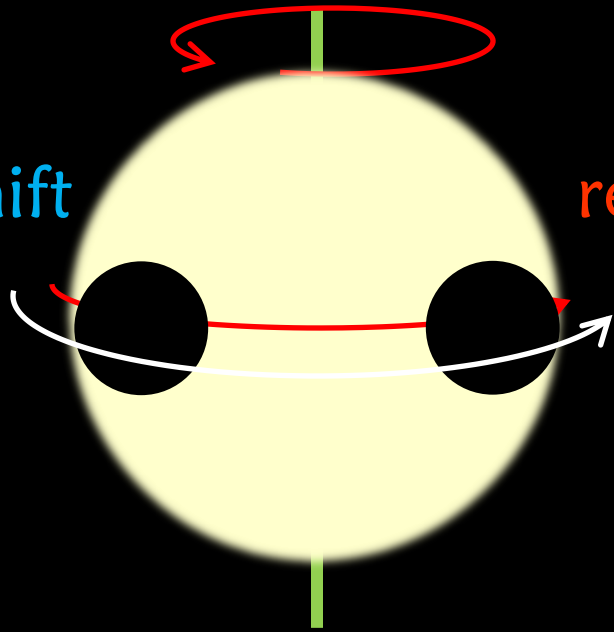
NEAR mission NASA

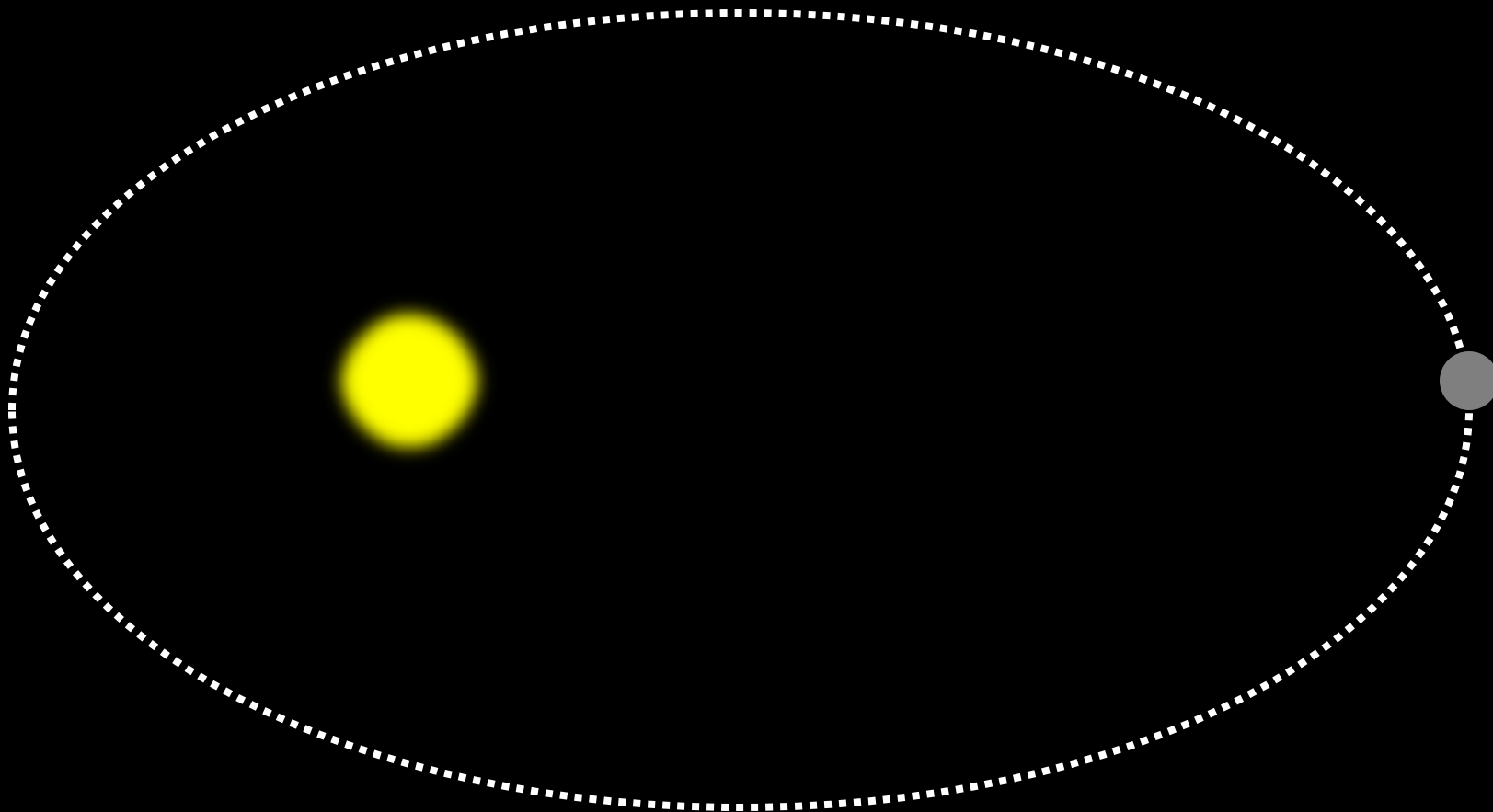


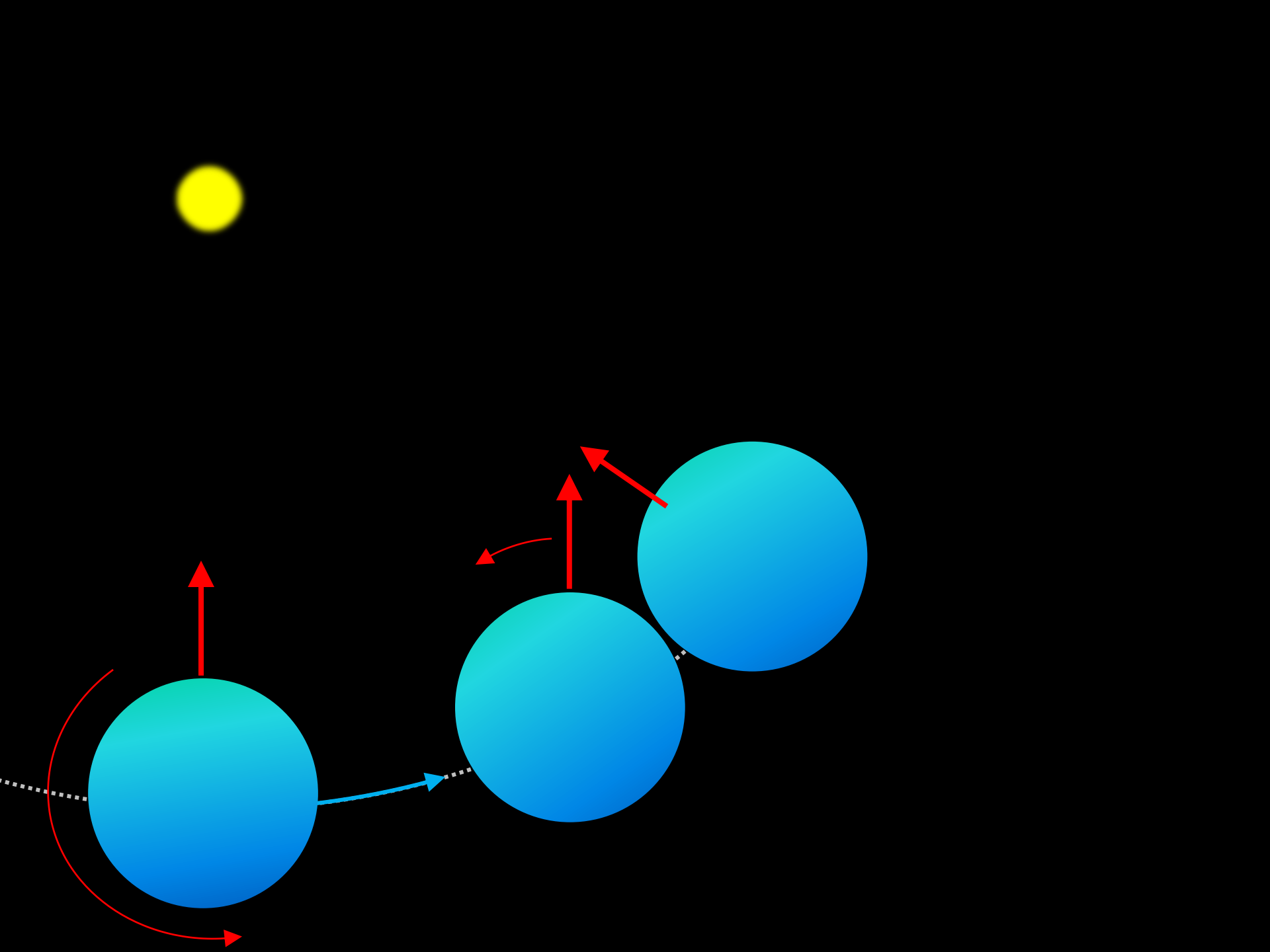


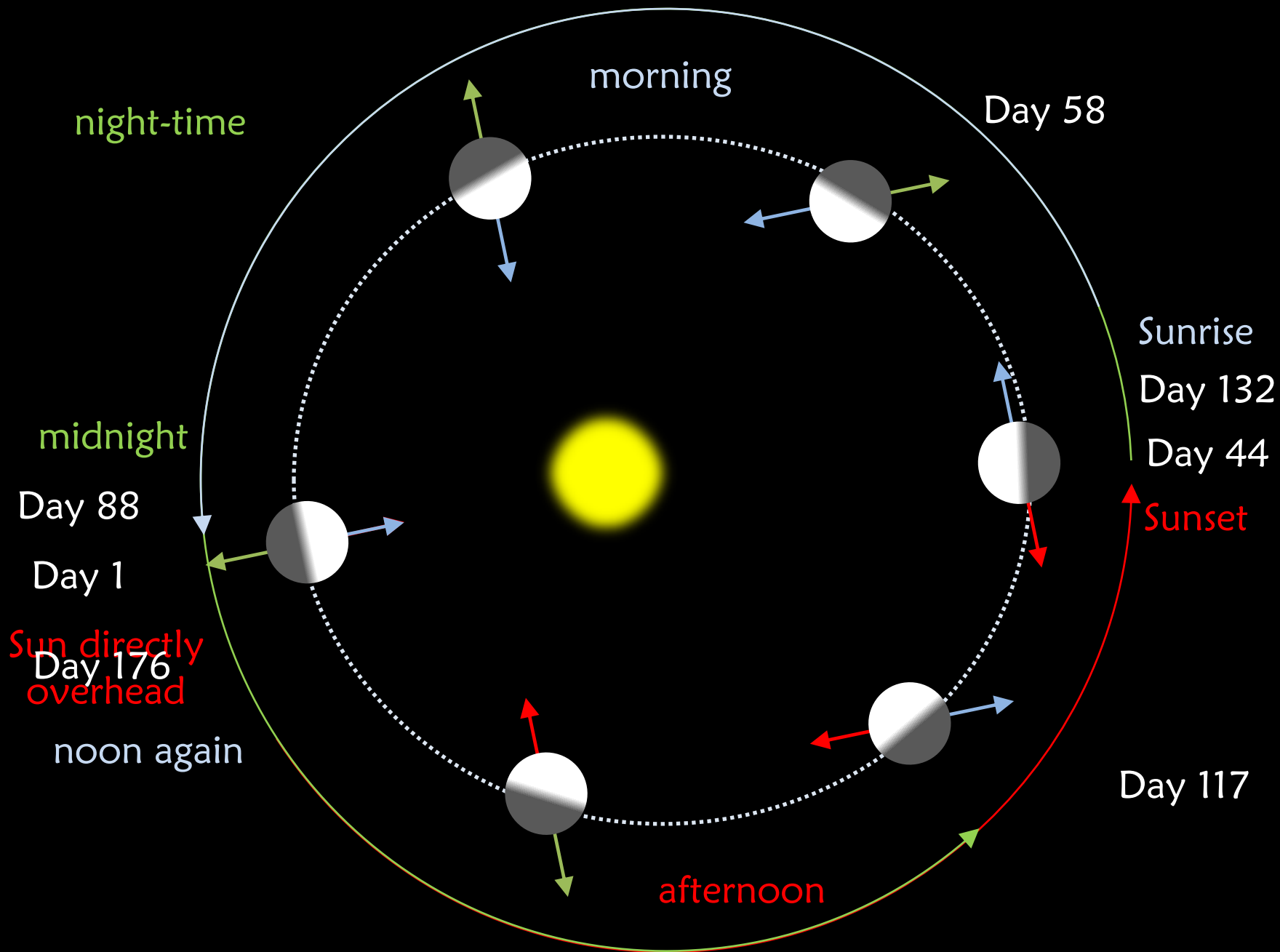
blueshift

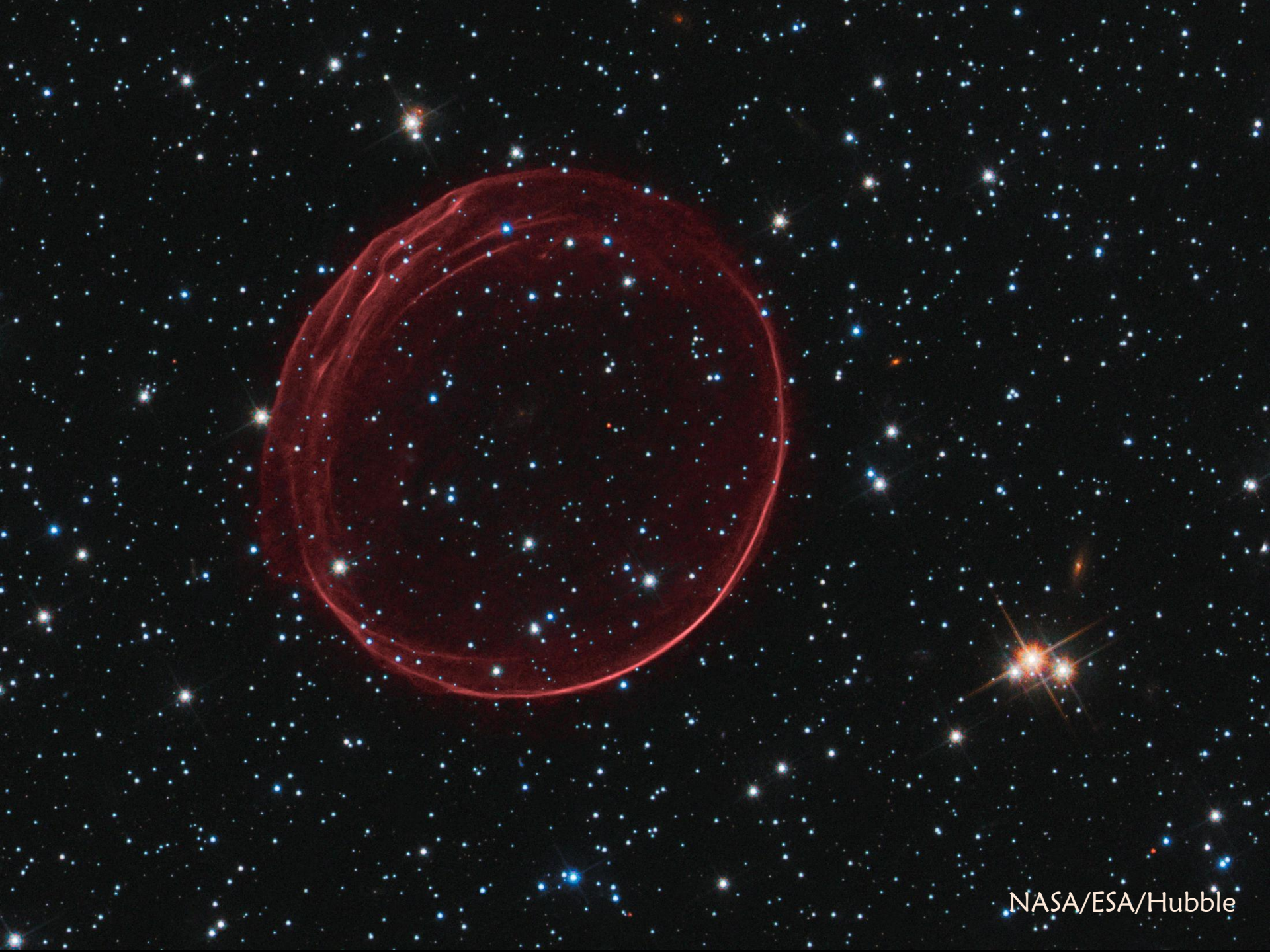
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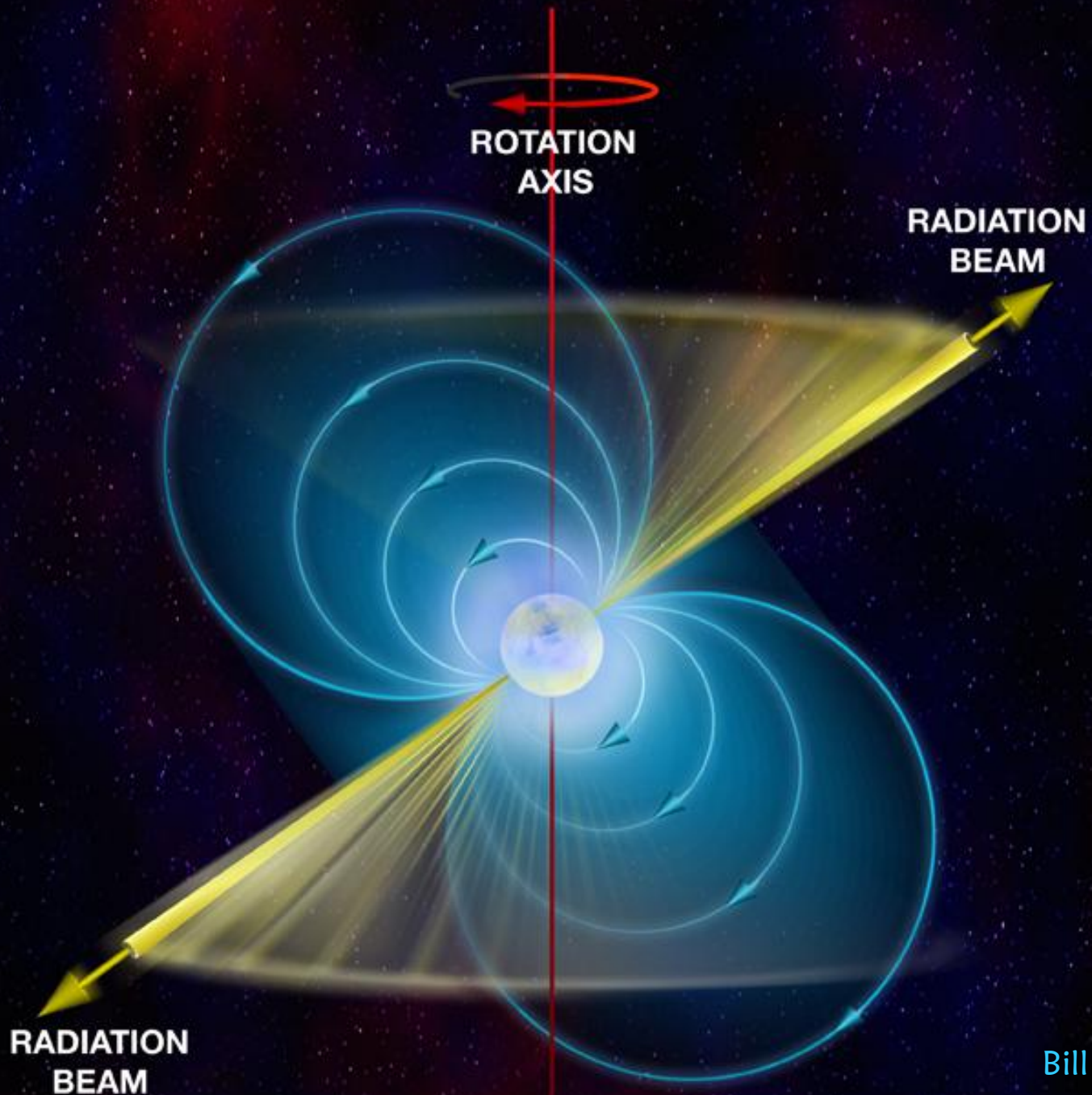




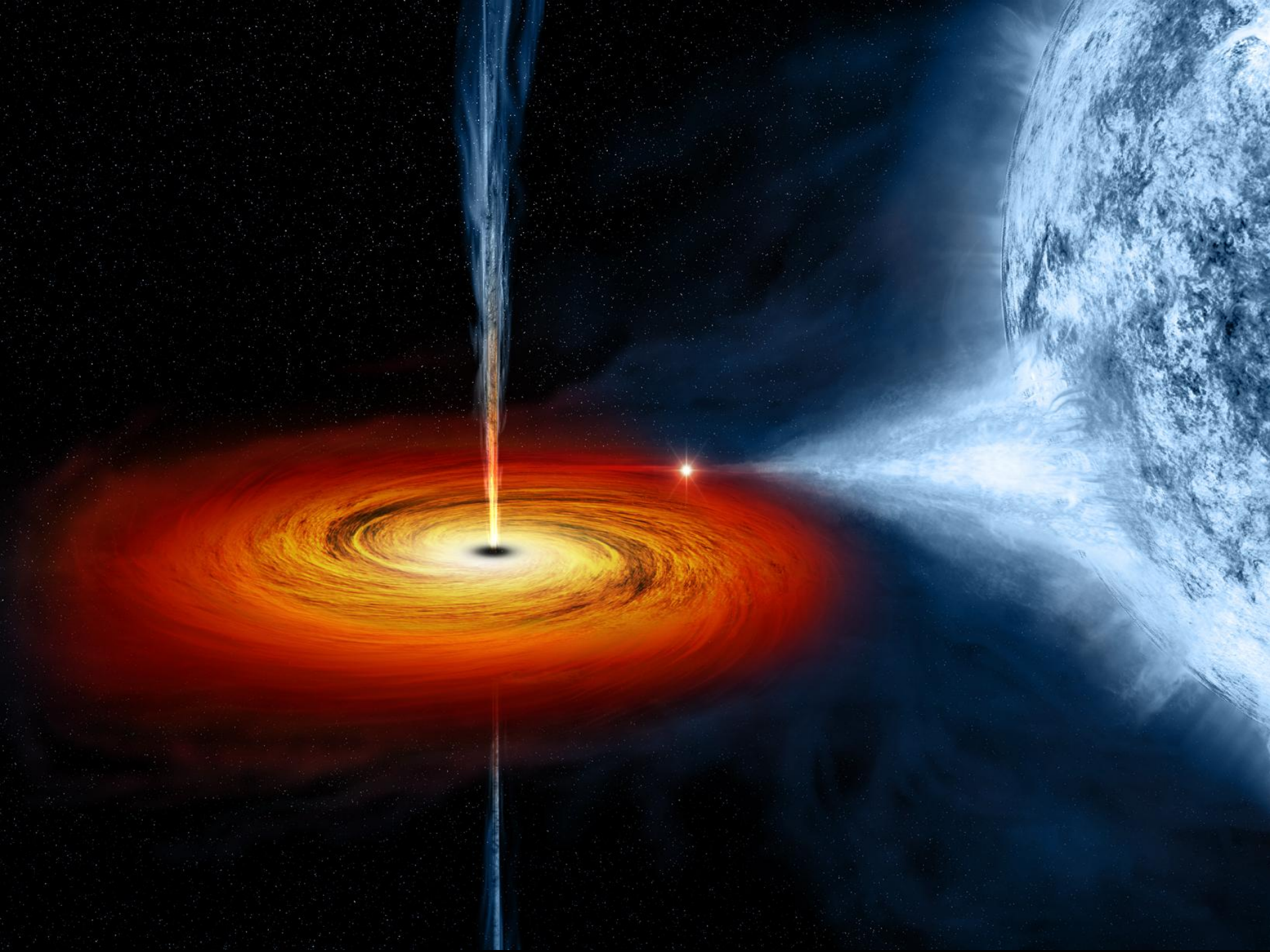


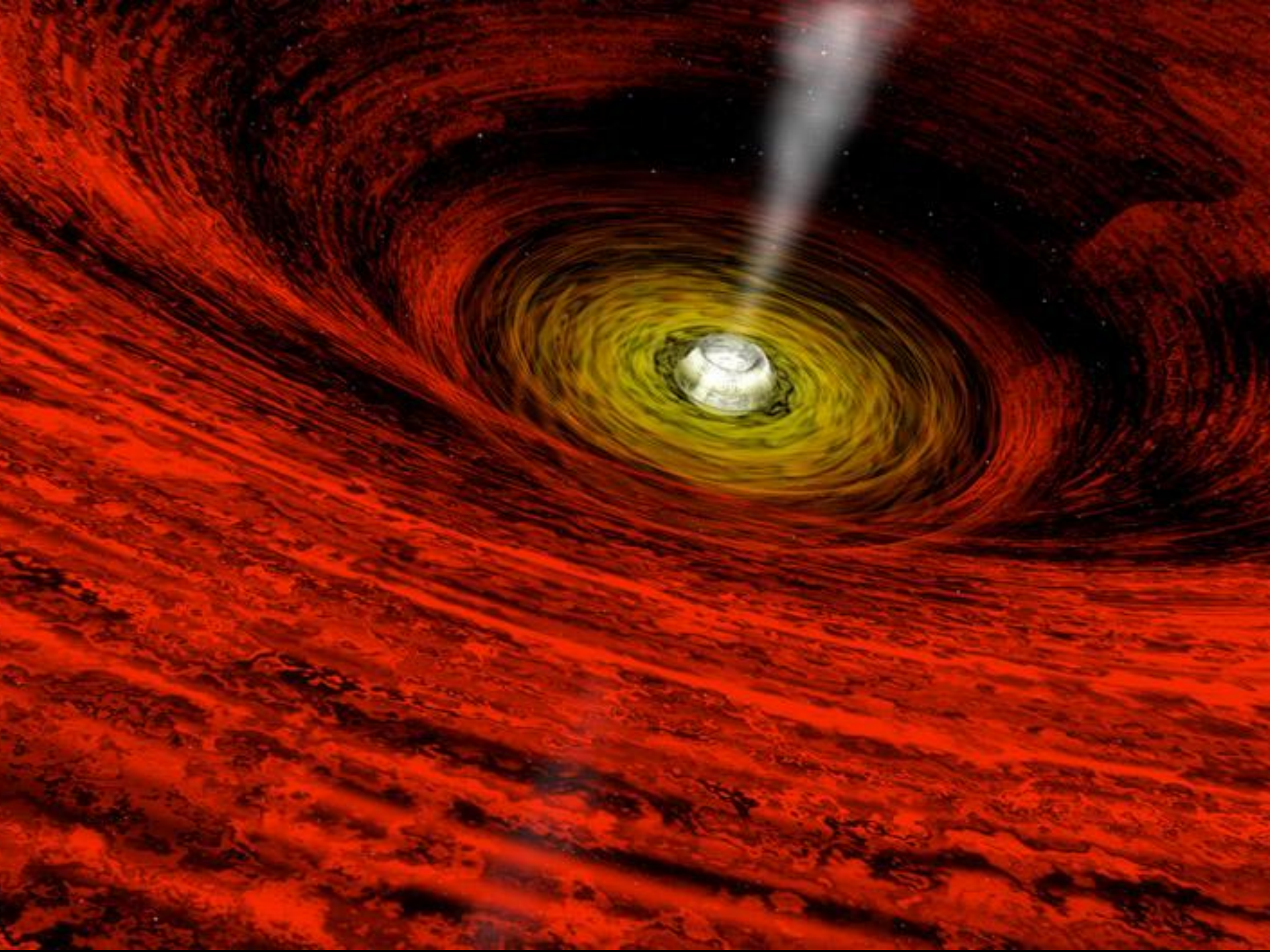


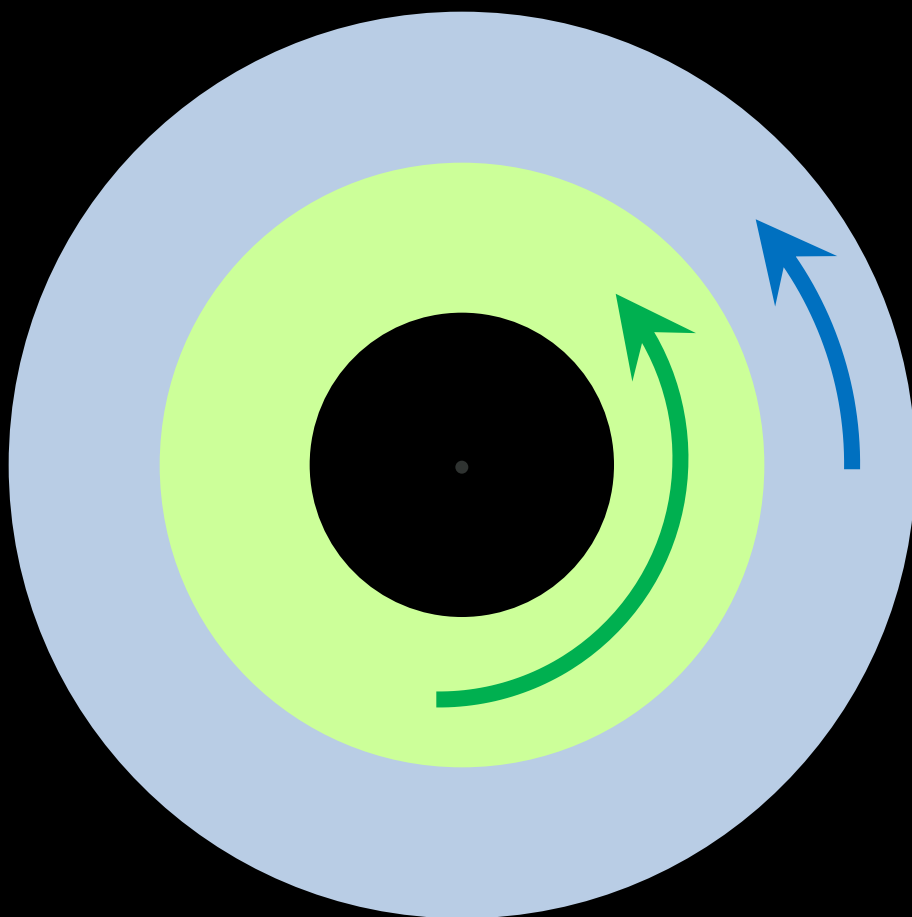


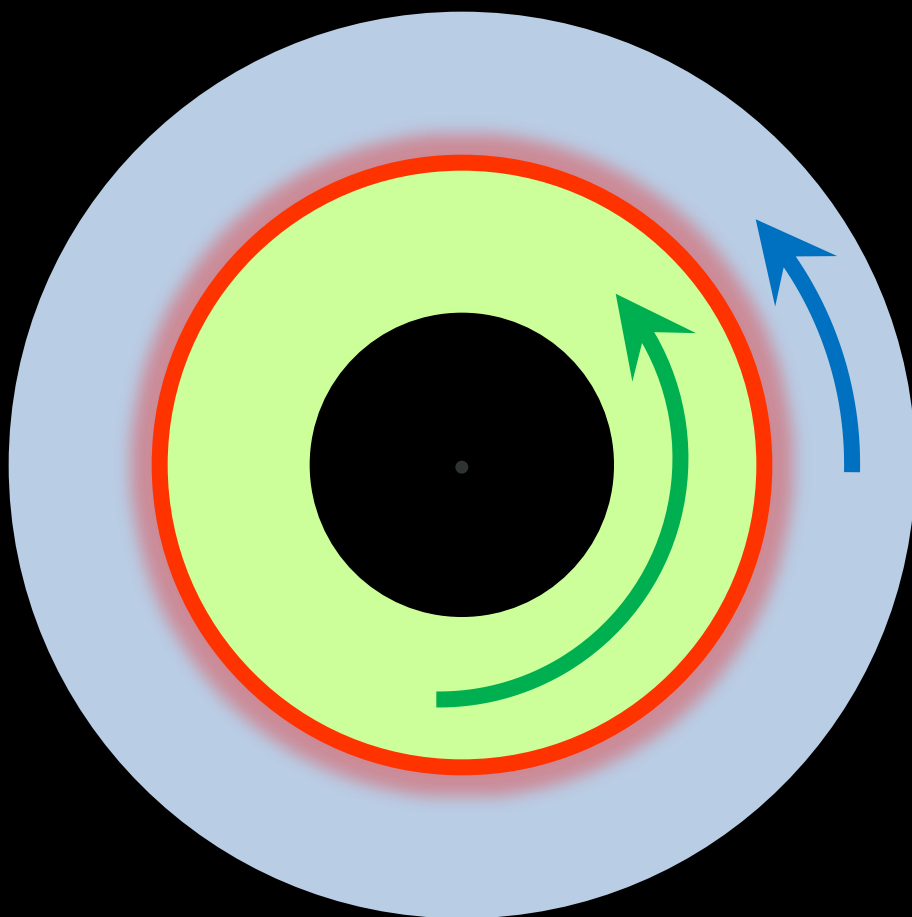


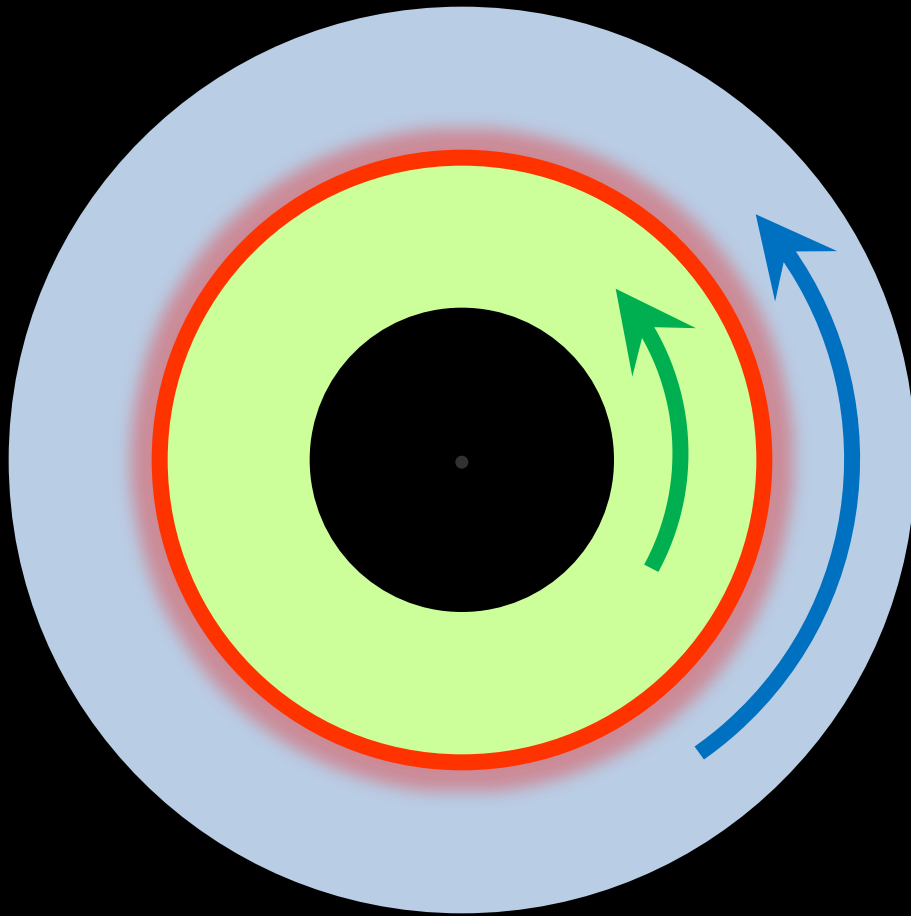
Bill Saxton, NRAO/AUI/NSF



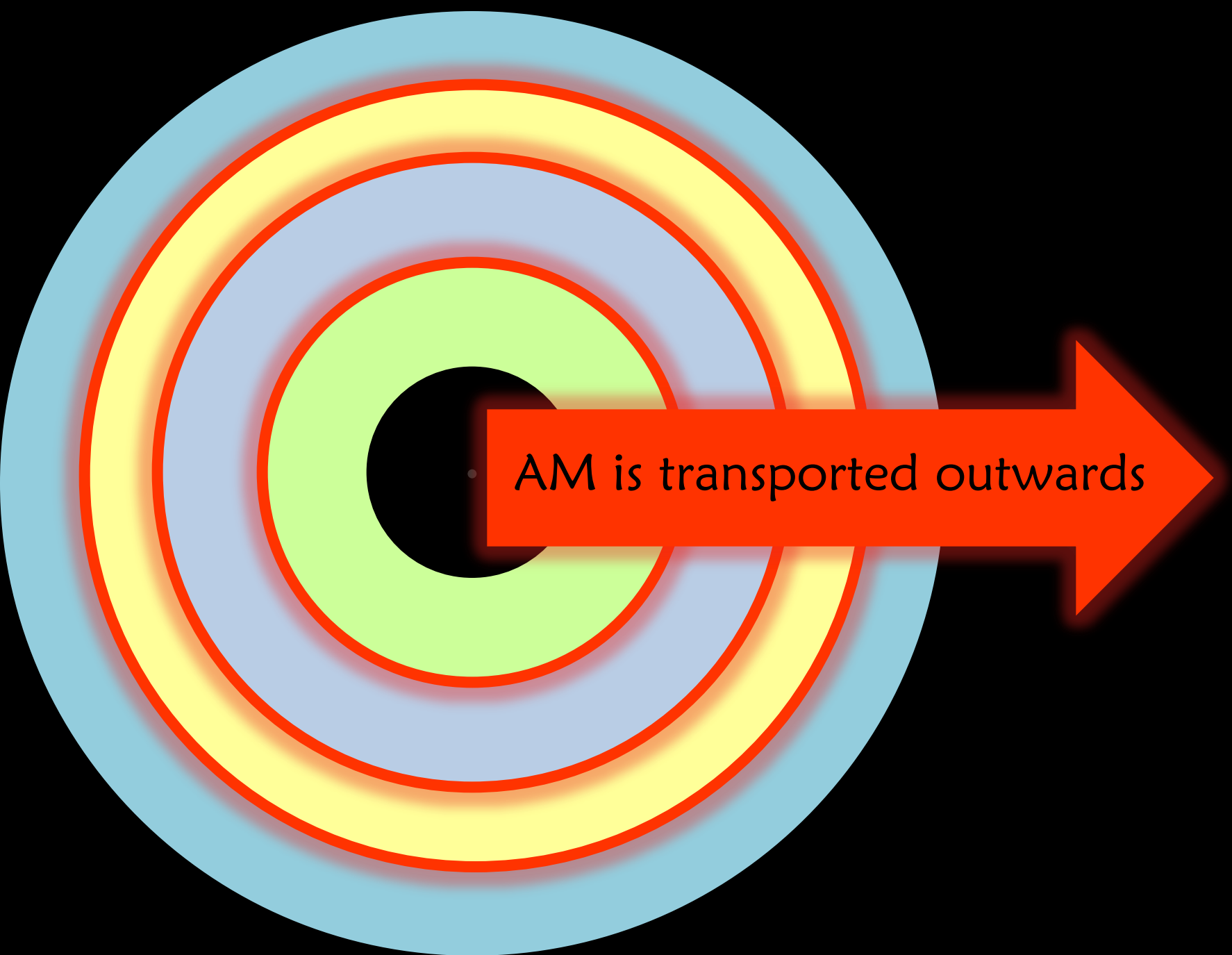




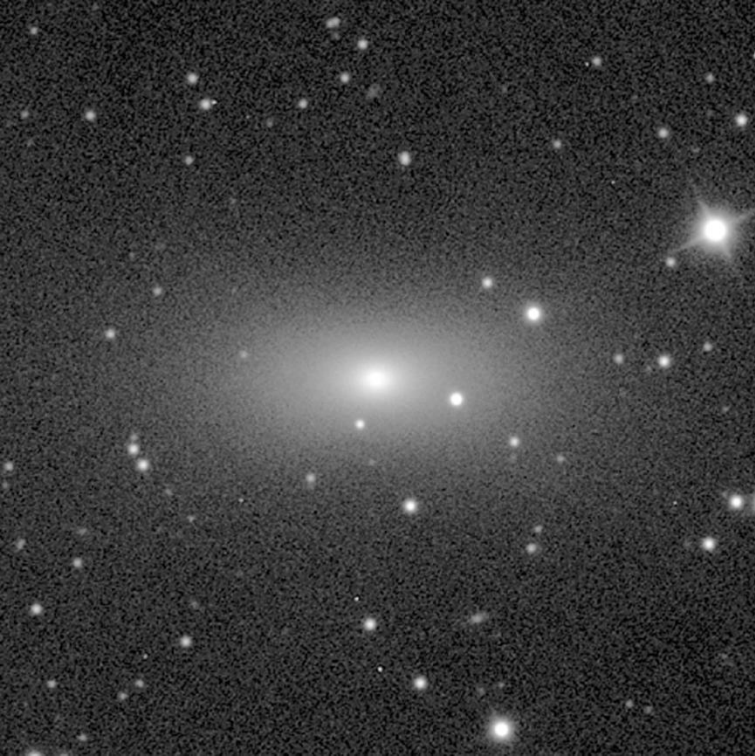




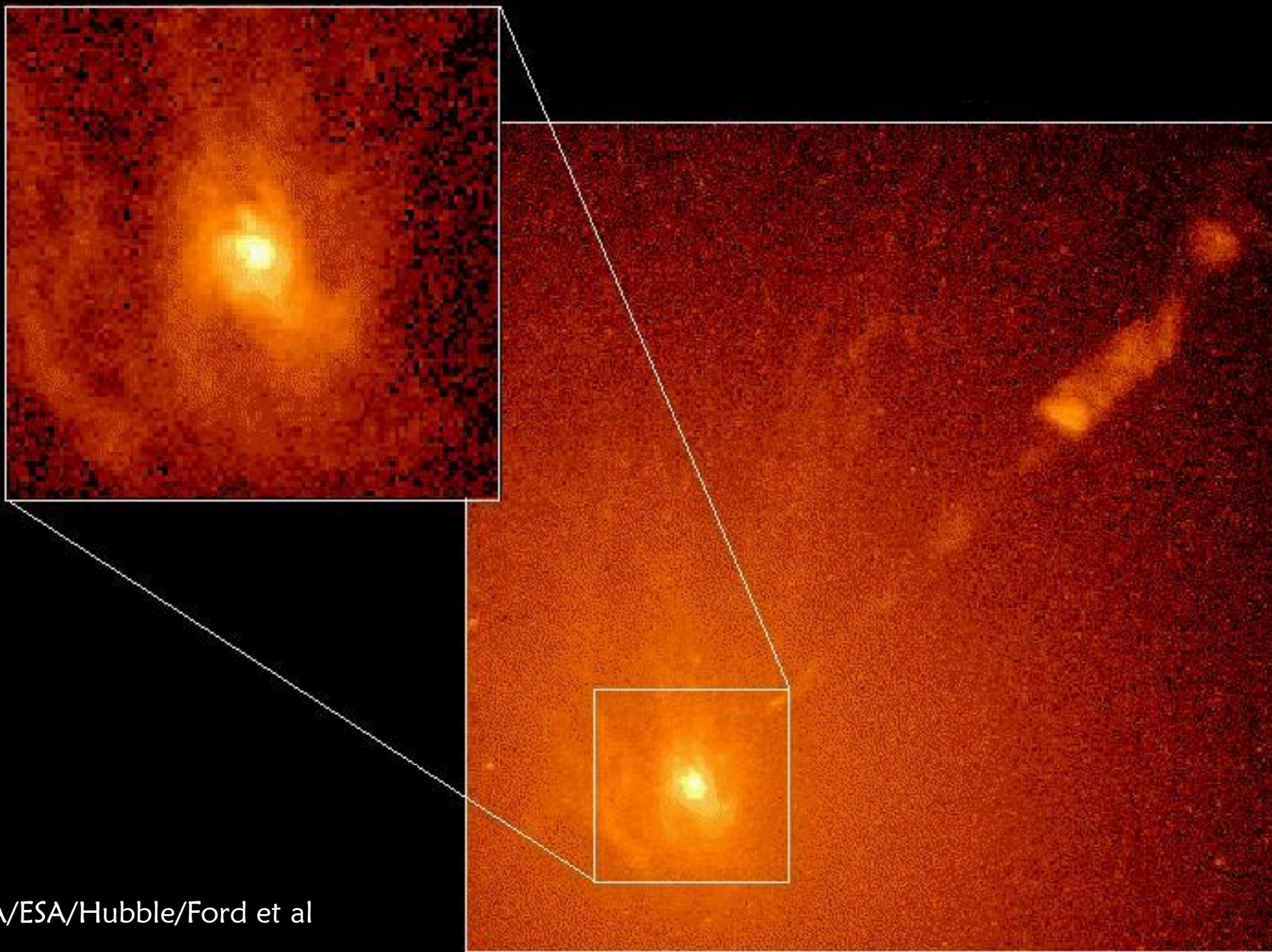
mass  $\times$  rotational rate  $\times$  radius<sup>2</sup>



AM is transported outwards

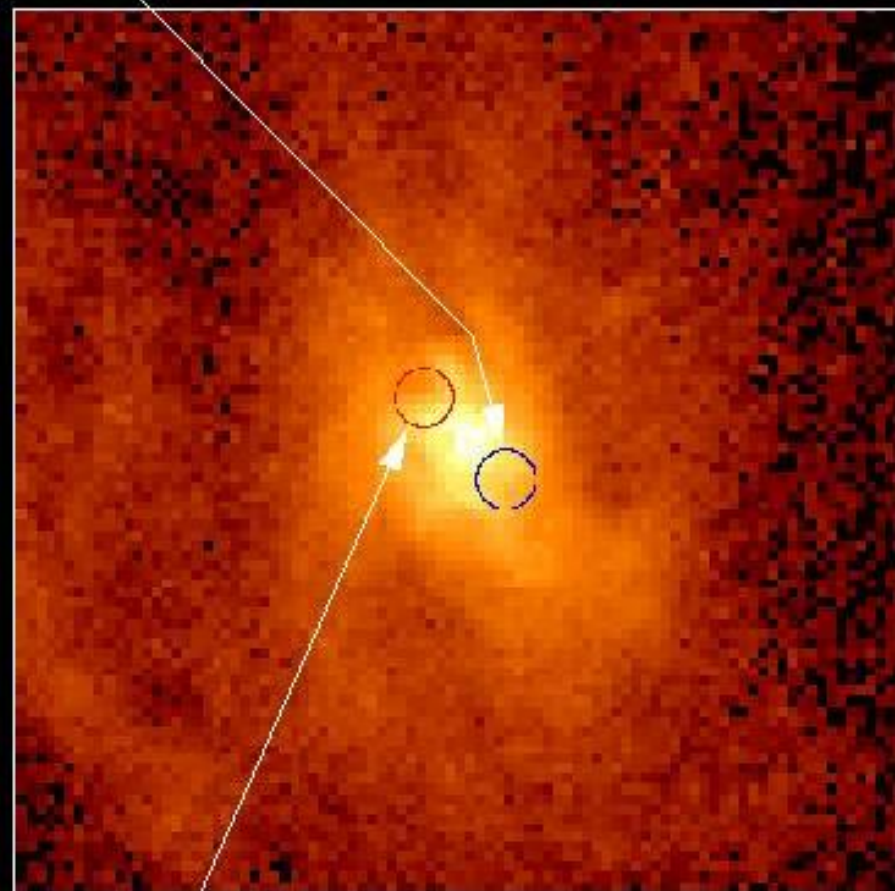




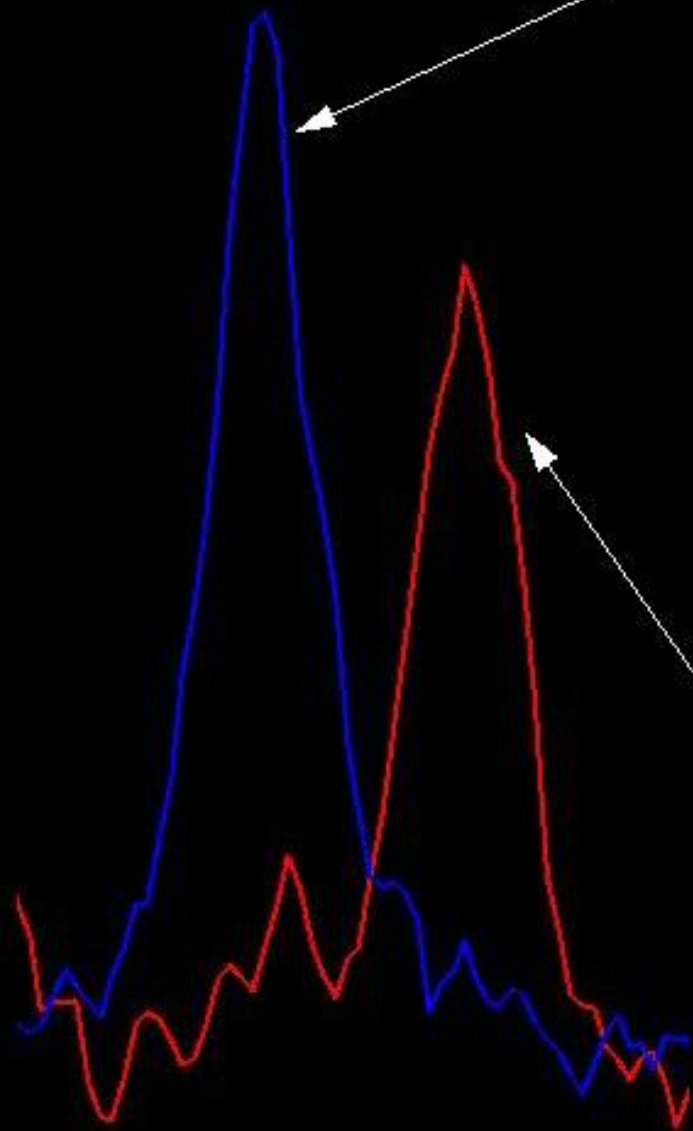


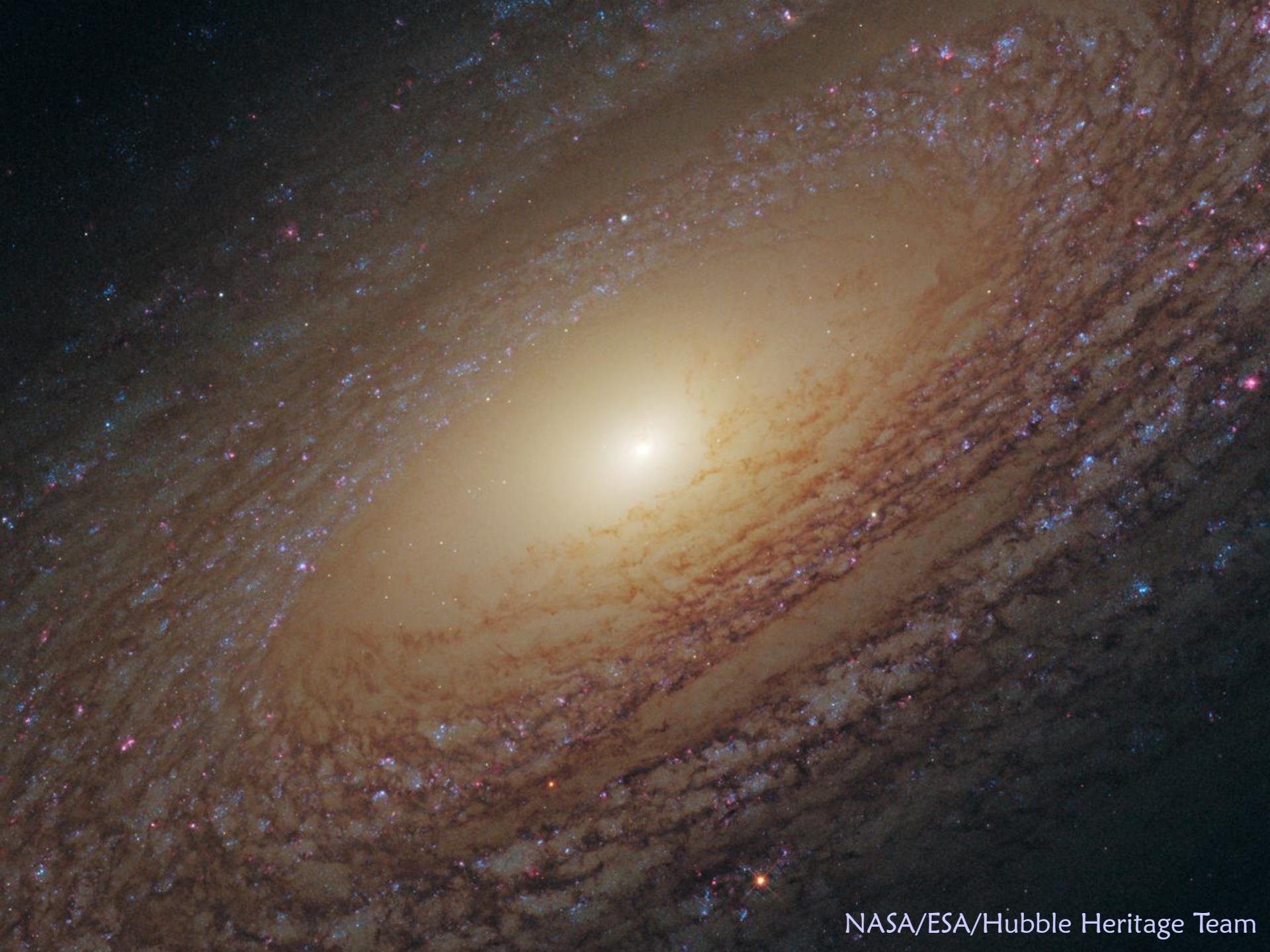
NASA/ESA/Hubble/Ford et al

Approaching



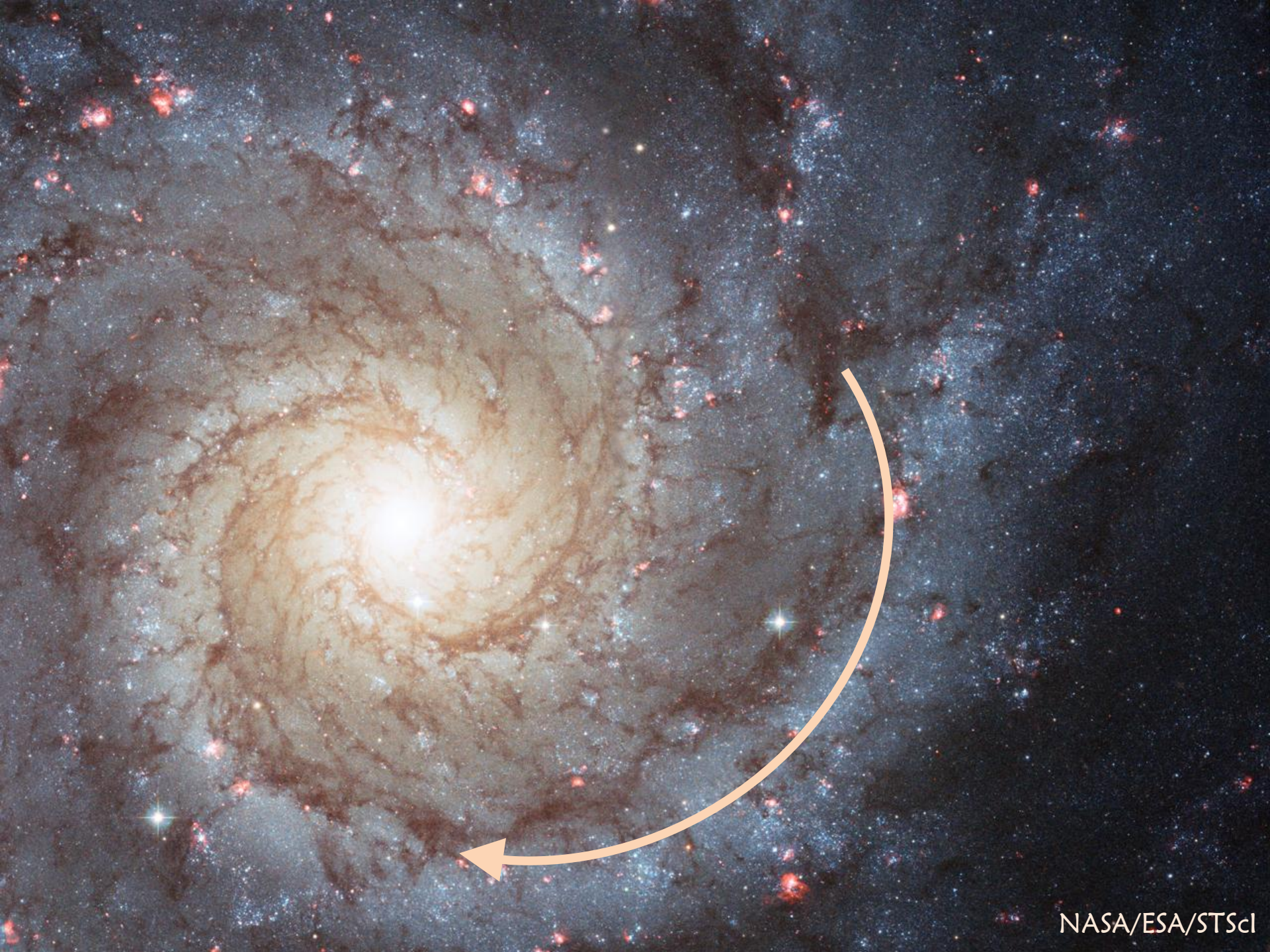
Receding

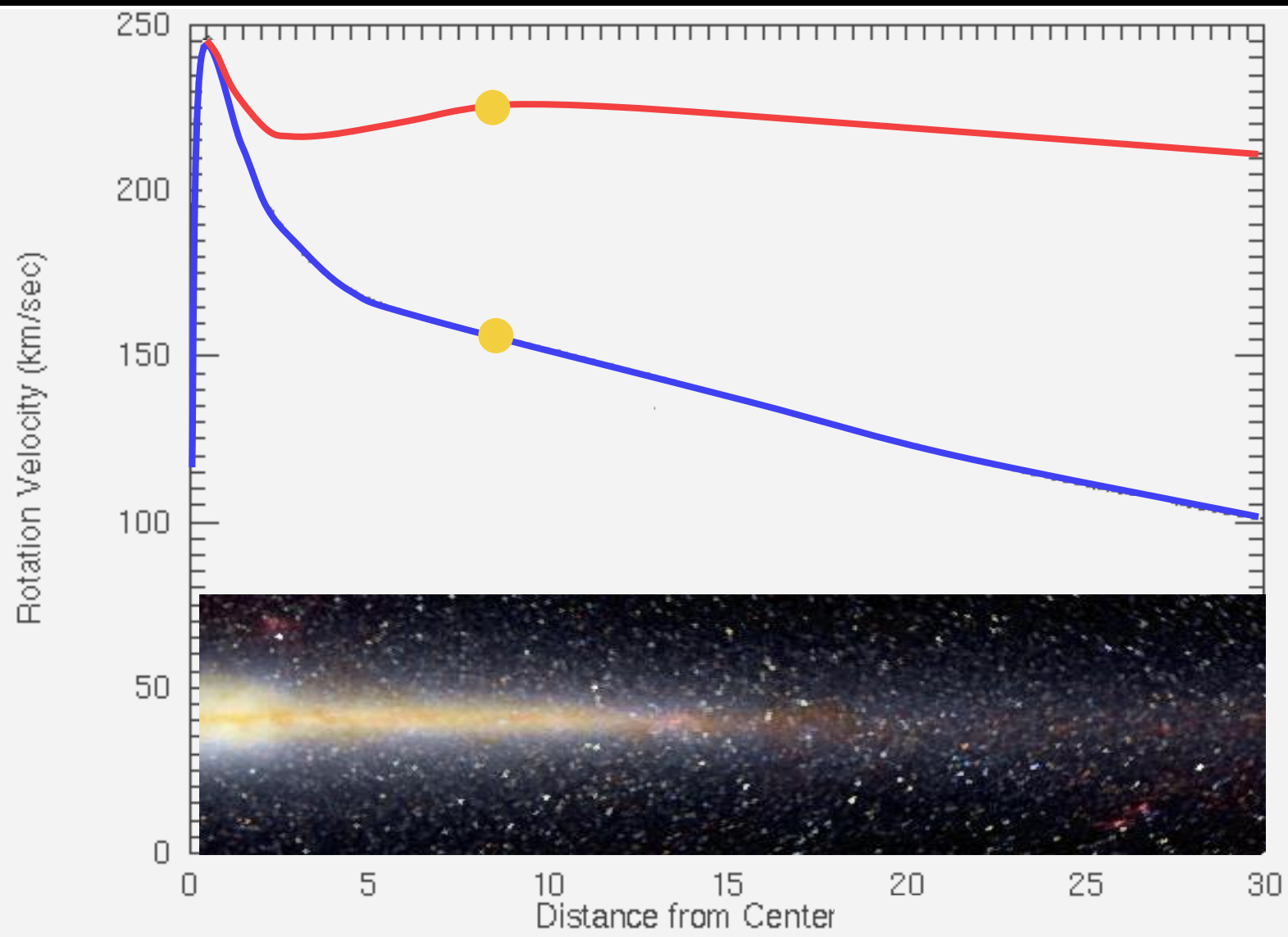




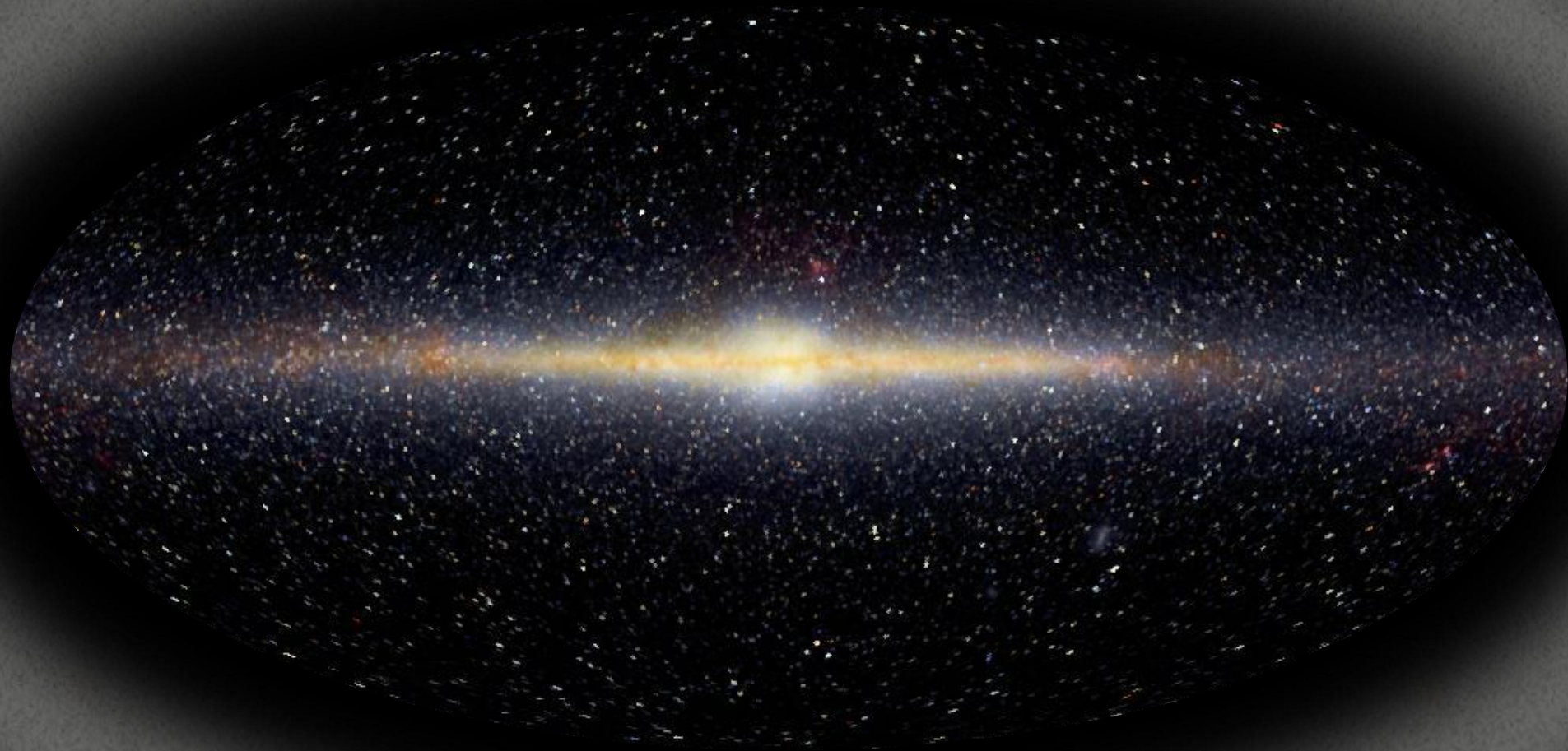
NASA/ESA/Hubble Heritage Team







# Dark Matter



# CLUSTERS OF GALAXIES

WEDNESDAY 7<sup>TH</sup> MARCH 2012 1PM

